UNIVERSITY^{OF} BIRMINGHAM

University of Birmingham Research at Birmingham

The rise of the ruling reptiles and ecosystem recovery from the Permo-Triassic mass extinction

Ezcurra, Martin; Butler, Richard

DOI:

10.1098/rspb.2018.0361

License

Other (please specify with Rights Statement)

Document Version
Peer reviewed version

Citation for published version (Harvard):

Ezcurra, M & Butler, R 2018, 'The rise of the ruling reptiles and ecosystem recovery from the Permo-Triassic mass extinction', *Proceedings of the Royal Society B: Biological Sciences*, vol. 285, no. 1880. https://doi.org/10.1098/rspb.2018.0361

Link to publication on Research at Birmingham portal

Publisher Rights Statement:

Checked for eligibility: 25/05/2018

The rise of the ruling reptiles and ecosystem recovery from the Permo-Triassic mass extinction, Martín D. Ezcurra, Richard J. Butler, Proc. R. Soc. B 2018 285 20180361; DOI: 10.1098/rspb.2018.0361. Published 13 June 2018

General rights

Unless a licence is specified above, all rights (including copyright and moral rights) in this document are retained by the authors and/or the copyright holders. The express permission of the copyright holder must be obtained for any use of this material other than for purposes permitted by law.

- •Users may freely distribute the URL that is used to identify this publication.
- •Users may download and/or print one copy of the publication from the University of Birmingham research portal for the purpose of private study or non-commercial research.
- •User may use extracts from the document in line with the concept of 'fair dealing' under the Copyright, Designs and Patents Act 1988 (?)
- •Users may not further distribute the material nor use it for the purposes of commercial gain.

Where a licence is displayed above, please note the terms and conditions of the licence govern your use of this document.

When citing, please reference the published version.

Take down policy

While the University of Birmingham exercises care and attention in making items available there are rare occasions when an item has been uploaded in error or has been deemed to be commercially or otherwise sensitive.

If you believe that this is the case for this document, please contact UBIRA@lists.bham.ac.uk providing details and we will remove access to the work immediately and investigate.

Download date: 09. Apr. 2024

The rise of the ruling reptiles and ecosystem recovery from the Permian-

_	2	Trioggio magg	oxtination
7	7	Triassic mass	extinction

3

1

4 Martín D. Ezcurra^{1,2}* and Richard J. Butler²*

5

- 6 ¹Sección Paleontología de Vertebrados, CONICET-Museo Argentino de Ciencias Naturales,
- 7 Ángel Gallardo 470 C1405DJR, Buenos Aires, Argentina.
- 8 ²School of Geography, Earth and Environmental Sciences, University of Birmingham,
- 9 Edgbaston, Birmingham B15 2TT, UK.
- *Correspondence to: martindezcurra@yahoo.com.ar, r.butler.1@bham.ac.uk

11

12

Abstract

- One of the key faunal transitions in Earth history occurred after the Permo-Triassic mass
- extinction (ca. 252.2 Ma), when the previously obscure archosauromorphs (which include
- 15 crocodylians, dinosaurs, and birds) become the dominant terrestrial vertebrates. Here, we
- place all known middle Permian-early Late Triassic archosauromorph species into an explicit
- 17 phylogenetic context, and quantify biodiversity change through this interval. Our results
- indicate the following sequence of diversification: a morphologically conservative and
- 19 globally distributed post-extinction 'disaster fauna'; a major but cryptic and poorly sampled
- 20 phylogenetic diversification with significantly elevated evolutionary rates; and a marked
- 21 increase in species counts, abundance, and disparity contemporaneous with global ecosystem
- stabilisation some 5 million years after the extinction. This multiphase event transformed
- 23 global ecosystems, with far-reaching consequences for Mesozoic and modern faunas.
- 24 **Keywords:** adaptive radiation; biotic crisis; morphological disparity; evolutionary rates;
- 25 Diapsida; Archosauromorpha

27

28

29

30

31

32

33

34

35

36

37

38

39

40

41

42

43

44

45

46

47

48

49

50

1. Introduction

The devastating Permo-Triassic (PT) mass extinction (ca. 252.2 Ma) dramatically impacted and remodelled global ecosystems [1–3]. On land, one of the key faunal transitions in Earth history took place during and following this extinction. The Palaeozoic amniote fauna, including synapsid groups such as anomodonts and gorgonopsians and parareptiles such as pareiasaurs, were decimated and largely displaced during the earliest Mesozoic by the previously obscure archosauromorphs [4,5]. Archosauromorphs, which include the 'ruling reptiles' or archosaurs (crocodylians, pterosaurs, dinosaurs, and their descendants, birds) and their close relatives, dominated terrestrial ecosystems for most of the Mesozoic and remain highly abundant and diverse in the modern biota [6–8]. Archosauromorphs originated during the middle–late Permian [9] and underwent a major radiation during the Triassic [6,10]. In the 20 million years following the PT mass extinction, species counts for archosauromorphs increased (>100 valid species currently known) and the group achieved high morphological diversity, including highly specialised herbivores, large apex predators, marine predators, armoured crocodile-like forms, and gracile dinosaur precursors [6,10]. Despite this high diversity, scientific attention has mainly focused on the diversification of crown archosaurs (particularly bird-line archosaurs [6–8,10–13]), and the early diversification of archosauromorphs around the PT boundary has often been overlooked and little discussed (e.g. [14]). Thus, the patterns and processes of the ascendance of archosauromorphs to dominance by the Late Triassic are incompletely explored and poorly understood. Comprehensive macroevolutionary analysis of the dawn of the archosauromorph radiation has been hampered by the absence of a comprehensive, explicit phylogenetic framework for these early species. Here, we quantitatively document major patterns of early archosauromorph biodiversity change, using a new phylogenetic dataset that includes for the first time all 108

currently valid middle Permian—early Late Triassic species (electronic supplementary material). Our analyses of morphological disparity, observed species counts, phylogenetic diversity, and rates of phenotypic evolution are focused on the first 35 million years of archosauromorph evolution (ca. 269–233 Ma) (figure 1a). These analyses aim to quantitatively explore one of the most important evolutionary radiations of vertebrates in the fossil record and the evolutionary patterns resulting from the reshaping and recovery of ecosystems in the aftermath of the deadliest mass extinction in Earth history.

2. Materials and methods

(a) Taxon-character data matrix

The quantitative macroevolutionary analyses conducted here are based on the most comprehensive species-level phylogenetic dataset currently available for early archosauromorphs [10] and its subsequent modifications (electronic supplementary material). We expanded this discrete morphological character matrix with the addition of 27 independent terminals (see supplementary table 1), which resulted in a new dataset composed of 149 terminals and 688 characters. However, the full dimensions of this dataset are 689 characters and 151 terminals because character 119 was deactivated a priori and there are two additional taxonomic units representing the scorings of the complete hypodigms of *Archosaurus rossicus* (electronic supplementary material) and *Osmolskina czatkoviensis* (supplementary table 1). These two terminals are not completely independent from the terminals representing the holotypes of these two species. In addition, some scorings were modified from previous versions of this data set (electronic supplementary material).

(b) Phylogenetic analysis

Phylogenetic diversity and evolutionary rates calculations require explicit phylogenetic hypotheses [15,16]. As a result, the complete data matrix including all 149 sampled terminals (including the complete hypodigm of *Osmolskina czatkoviensis*; supplementary table 1) was analysed under equally weighted maximum parsimony using TNT 1.5 [17] in order to recover the required phylogenetic trees. The search strategy used a combination of the tree search algorithms Wagner trees, TBR branch swapping, sectorial searches, Ratchet (perturbation phase stopped after 20 substitutions), and Tree Fusing (5 rounds), and continued until the same minimum tree length was hit 100 times. The best trees obtained using this strategy were subjected to a final round of TBR branch swapping. Zero length branches in any of the recovered most parsimonious trees (MPTs) were collapsed and several characters were considered additive (electronic supplementary material).

(c) Time bins

The aim of our analyses is to explore the first 35 million years of the evolutionary history of Archosauromorpha, spanning the Permian origins of the group through to the appearance of archosauromorph-dominated ecosystems in the late Middle Triassic and earliest Late Triassic. We used five time bins in order to examine macroevolutionary patterns during this time span: middle–late Permian (~17.1 myr), Induan (1.0 myr), Olenekian (4.0 myr), Anisian (5.2 myr), and Ladinian–early Carnian (~9.0 myr) [18]. Despite the very short length of the Induan, this stage was maintained as a separate time bin in order to capture diversity changes that occurred in the immediate aftermath of the PT mass extinction.

(d) Temporal calibration of trees

The evolutionary rates analyses require time-calibrated trees. The trees were calibrated with the timePaleoPhy() function of the package paleotree for R [19] using the "mbl" calibration

[11,20], a minimum branch length of 0.1 myr, and a root age of 269.3 Ma based on the maximum bound estimated for the origin of Archosauromorpha [9] (figure 1*a*; supplementary figure 4). Sensitivity analyses to explore the effect that different temporal calibrations may have on the results of the evolutionary rate analyses were conducted using "mbl" calibrations with minimum branch lengths of 0.5 and 1.0 myr, and also using the "cal3" method [21] (electronic supplementary material).

106

107

108

109

110

111

112

113

114

115

116

117

118

119

120

121

122

123

124

100

101

102

103

104

105

(e) Morphological disparity analyses

Changes in morphological diversity (disparity) were quantified using the R package Claddis [16]. All non-archosauromorph species and archosauromorphs that occur in late Carnian or younger stratigraphic horizons were pruned before the disparity analyses, resulting in a final dataset of 112 terminals. Some terminals occur across two time bins because of uncertainty in the dating of the stratigraphic unit from which their fossils have been collected. These taxa were counted in both time bins in the disparity analyses (supplementary tables 2, 3). A sensitivity analysis pruning these terminals with stratigraphic uncertainty was conducted to evaluate the effect on the results (electronic supplementary material). Disparity curves were reconstructed using both Generalized Euclidean Distance (GED) and Maximum Observable Rescaled Distance (MORD) dissimilarity matrices (the two distance matrices recommended by Lloyd [16] for conducting disparity analyses based on discrete characters) generated from the taxon-character data matrix after the a priori pruning of non-archosauromorphs and those archosauromorph taxa stratigraphically younger than early Carnian (electronic supplementary material). These dissimilarity matrices were used to calculate weighted mean pairwise dissimilarity (WMPD) as a disparity metric. Statistical significance between the disparity metrics for each time bin was assessed through 95% confidence intervals calculated from 1,000 bootstrap replicates of the original taxon-character data matrix and a recalculation of

the dissimilarity matrices and disparity metrics. Morphospace bivariate plots were generated for the entire data set and each time bin based on the results of a Principal Coordinate Analysis performed on the GED dissimilarity matrix. An additional disparity analysis using the same archosauromorph sampling as Foth et al. [14] was conducted using the same protocol.

130

131

132

133

134

135

136

137

138

139

140

141

142

143

144

145

146

147

148

125

126

127

128

129

(f) Phenotypic evolutionary rates analyses

Ancestral character-states were reconstructed with the package claddis [16] using maximum likelihood in order to infer significant departures from equal rates of character evolution [22]. The phylogenetic analysis of the dataset compiled here recovered more than 10,000 MPTs. Therefore, in order to reduce computational time we used a random sample of 100 of these trees for the main evolutionary rate analyses (figure 2a). Non-archosauromorph terminals were pruned, but archosauromorph terminals stratigraphically younger than the early Carnian were retained because of the effects that the ghost lineages that they generate may have on older time bins (electronic supplementary material). All 100 subsampled trees were temporally calibrated using the protocol described above. The evolutionary rate analysis was conducted using the function DiscreteCharacterRate() {Claddis}, setting an alpha of 0.01 (supplementary figure 8). An alpha of 0.01 was preferred because, as stated by Lloyd [16], there is generally a high heterogeneity of rates within data sets. A reduction in the alpha value therefore represents a conservative approach to reduce the number of significant values. Confidence intervals for each time bin were calculated using the function plotMeanTimeseries (), written by Close et al. [23], in order to test for the presence of significant rate differences in the early evolutionary history of Archosauromorpha (table 1). Sensitivity analyses using alternative tree calibrations were conducted using 10 trees for each

"mbl" calibration and the 60 trees generated by the "cal3" method (electronic supplementary material).

(g) Time series comparisons

Some of the macroevolutionary metrics calculated here may be correlated with one another and should not be considered as independent. To test this, we made statistical comparisons between observed species counts, phylogenetic diversity, specimen-level abundance data (i.e. number of individuals), and number of archosauromorph-bearing formations (as a metric of fossil record sampling). To compare these time series we used generalized least-squares regression (GLS) with a first order autoregressive model (corARMA) fitted to the data using the function gls() in the R package nlme v. 3.1–137 [24]. GLS reduces the chance of overestimating statistical significance of regression lines due to serial correlation. Time series were not log-transformed prior to analysis, as none were non-normally distributed (Shapiro-Wilk tests p>0.1). We calculated likelihood-ratio based pseudo-R² values using the function r.squaredLR() of the R package MuMIn [25]..

3. Results

Our results show a significant decrease in morphological disparity (using a Maximum Observable Rescaled Distance dissimilarity matrix, MORDdm) or a non-significant change (using a Generalized Euclidean Distance dissimilarity matrix, GEDdm) from the middle–late Permian to the earliest Triassic (Induan). Subsequently, a dramatic, significant increase occurs in the Olenekian (using MORDdm) or Anisian (using GEDdm) and high disparity levels are maintained in the Ladinian–early Carnian (figures 2b, 3; table 1). Evolutionary rates are significantly higher during the Olenekian—and in several topologies also during the Induan—than in other intervals (figure 2a and table 1), coincident with a peak in

phylogenetic diversity (figure 1b). This peak in phylogenetic diversity results from a number of phylogenetically deeply nested groups occurring in this interval, such as ctenosauriscids, which imply numerous ghost lineages [12] (figure 1a). Several of these lineages are identified as having significantly high evolutionary rates (e.g. supplementary figure 8). By contrast, significantly lower evolutionary rates are recovered for the Ladinian—early Carnian in all analyses (figure 2a) and also during the middle—late Permian using "mbl" calibrations (electronic supplementary material).

The observed or 'raw' species count of Induan archosauromorphs is at least double that recorded for the middle–late Permian, and observed species count increases only slightly during the Olenekian, but shows substantial increases into the Middle Triassic (figure 1*b*). Observed abundance data shows a pattern consistent with that for observed species count, with only very slight increases through the middle–late Permian to Olenekian time span followed by a remarkable increase in the Anisian (figure 1*b*). However, the time series of observed species count, number of individuals, and geological sampling (numbers of rock units in which archosauromorphs occur) are not significantly different to each other (p<0.05; pseudo-R²>0.85), which might reflect either a sampling bias or an increase of archosauromorph abundance in their ecosystems. Conversely, estimated phylogenetic diversity is not correlated with sampling estimates or abundance (p>0.15; R²<0.35) (supplementary table 6).

4. Discussion

Our analyses support a multiphase model of early archosauromorph diversification, largely in response to the events of the PT mass extinction. Archosauromorphs most likely originated in the middle Permian, and underwent a substantial phylogenetic diversification and dispersed across Pangea [9,26]. However, disparity remained low, and low fossil abundance (figures

1b, 2b, 3b, 3c) suggests either that archosauromorphs remained very minor components of

199

200

201

202

203

204

205

206

207

208

209

210

211

212

213

214

215

216

217

218

219

220

221

222

223

terrestrial ecosystems, or that this diversification took place in geographic regions or environments that remain poorly sampled. Many major lineages of archosauromorphs are inferred to have passed through the PT boundary and the group may have been comparatively little affected by the extinction event [10] (figure 1a). The Induan, immediately after the extinction, saw a substantial increase in archosauromorph abundance and a high observed species count relative to the length of the time bin, characterised by a low disparity (figure 2b), globally distributed archosauromorph 'disaster fauna' dominated by proterosuchids and a number of morphologically similar lineages (e.g. *Prolacerta*) [27] (figure 3b). This disaster fauna was apparently short-lived: in South Africa, *Proterosuchus* occurs only between 5–14 metres above the PT boundary [28]. Similar patterns have been documented for the synapsid Lystrosaurus following the PT extinction [29], and earliest Triassic tetrapod assemblages on land appear in general to have been highly uneven and dominated by a few highly abundant or diverse taxa [30,31]. Major perturbations in the global carbon cycle, referred to as 'chaotic carbon cycling', have been documented through the Early Triassic (Induan and Olenekian) [32,33] (figure 1c). These perturbations have been suggested to reflect either successive short-term greenhouse crises and rapid environmental change or boom-bust cycles of ecosystem instability [30,33,34]. This interval of instability coincides with generally elevated global temperatures that would have limited diversity in equatorial regions and a well-known gap in the coal record that reflects lowered plant productivity and diversity [34,35]. Our data suggest that archosauromorphs underwent a major phylogenetic diversification in the Olenekian (1–5

million years [myr] after the extinction), characterised by significantly elevated evolutionary

rhynchosaurs, archosaurs, erythrosuchids, and tanystropheids (figure 1a). The fossil record

rates (figure 2a), with the origins or initial diversification of major clades such as

225

226

227

228

229

230

231

232

233

234

235

236

237

238

239

240

241

242

243

244

245

246

247

248

shows that mass extinctions promote adaptive radiations in surviving, often previously marginal, clades because of the disappearance of species or entire lineages opening new vacancies in ecological space [36,37]. Thus, this general pattern suggests that the diversification of archosauromorphs was a response to vacant ecological space following the PT extinction, and the subsequent disappearance of the short-lived post-PT disaster fauna. However, observed species count and abundance remained low in the Olenekian, and similar to those of the Induan (figures 1b, 2b, 3b, 3c). As such, this major phylogenetic and probable morphological diversification in the Olenekian is at present largely cryptic and very incompletely sampled, potentially reflecting the very low abundances of individual archosauromorph species in the highly uneven and unstable Early Triassic ecosystems (figure 1b), as well as the limited geographical range over which known Olenekian tetrapod fossils occur [35]. The Anisian (5–10 myr after the extinction) is characterised by marked increases in observed species count, abundance, and disparity among archosauromorphs (figures 1b, 2b, 3d), as well as substantial increases in maximum body size [38]. An increased ecomorphological disparity during the Anisian matches previous results based on geometric morphometrics of archosauromorph skulls [14] (electronic supplementary material) and is documented in the skeletal fossil record by the appearance of large hypercarnivores, bizarre and highly specialised herbivores, long-necked marine predators, and gracile and agile dinosauromorphs [6,10]. This coincides with the end of the interval of intense carbon perturbations, a global cooling event, and the return of conifer-dominated forests [34], suggesting the recovery and stabilization of global ecosystems [30]. This stabilisation may have acted as an extrinsic factor that promoted increases in abundance among archosauromorph lineages as community evenness recovered, with a previously largely cryptic diversification becoming better sampled in the fossil record as a result. Similar

patterns are observed among marine tetrapods, with the first sauropterygians and ichthyosauromorphs being documented close to the Olenekian-Anisian boundary [39], but likely reflecting a temporally somewhat deeper period of currently unsampled phylogenetic diversification [40].

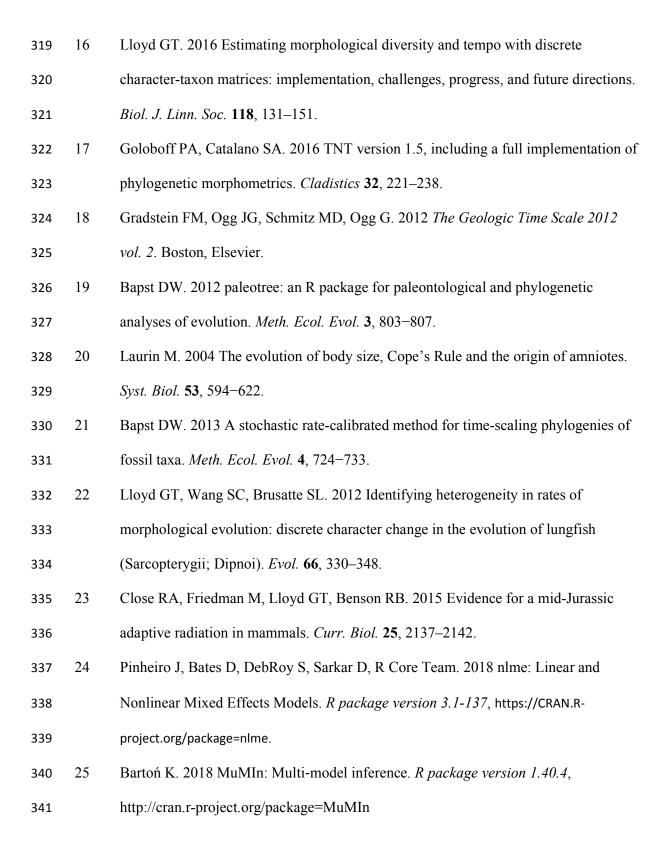
Our analyses of archosauromorph biodiversity change around the PT boundary support a diversity-first model of evolution, in which a rapid speciation of similar disaster taxa filled ecospace, followed by more steady adaptive evolution into new sectors of morphospace as ecosystems and community interactions stabilized (figure 3) [3]. A similar evolutionary pattern has been reported among dicynodonts in terrestrial ecosystems in the aftermath of the PT mass extinction [41], and has also been documented in fossil marine animals [42], including graptoloids [43] and ammonoids [44] during the Ordovician and PT biotic crises, respectively. More detailed work on other taxonomic groups is needed to establish if this pattern characterises other terrestrial clades and extinction events.

The establishment of high abundance, ecomorphological diversity, and observed species counts and phylogenetic diversity of archosauromorphs by the Middle Triassic paved the way for the ongoing diversification of the group (including the origins of dinosaurs, crocodylomorphs, and pterosaurs) in the Late Triassic, and their dominance of terrestrial ecosystems for the next 170 million years. Our results show the fundamental role of the PT mass extinction and its aftermath in reshaping terrestrial ecosystems, and its far-reaching impact on the faunas of the Mesozoic and modern world.

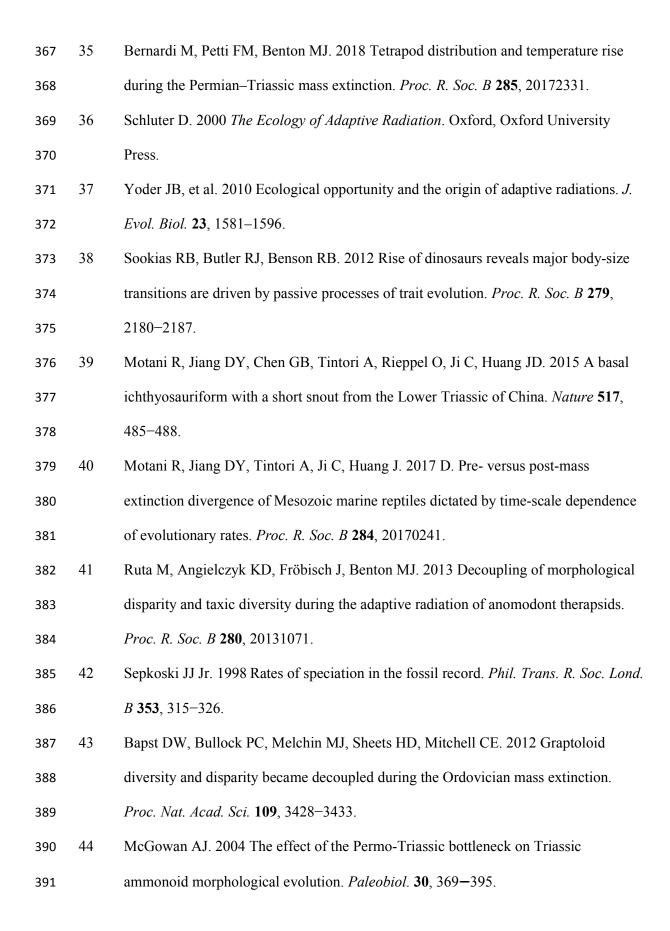
Data accessibility. Species occurrence data, R scripts, data matrices, and tree files are available as online electronic supplementary material.

272 **Authors' contributions.** M.D.E. and R.J.B. designed the research project, conducted the 273 analyses, and contributed to the text of the manuscript; M.D.E. scored most terminals and 274 made the figures. 275 **Competing interests.** We declare we have no competing interests. Funding. This research was supported by the DFG Emmy Noether Programme (BU 2587/3-1 276 277 to RJB), a Marie Curie Career Integration Grant (630123 ARCHOSAUR RISE to RJB), and a 278 National Geographic Society Young Explorers Grant (9467-14 to MDE). 279 **Acknowledgements.** We thank Roger Close and David Button for their comments and help 280 with some of the analyses. We also thank the associate editor Erin Saupe, Stephen Brusatte, 281 and an anonymous reviewer for their comments, which improved the manuscript. 282 283 References 284 1 Raup DM. 1979 Size of the Permo-Triassic bottleneck and its evolutionary 285 implications. Science 206, 217–218. 286 2 Erwin DH. 1994 The Permo-Triassic extinction. *Nature* **367**, 231–236. 287 3 Chen Z-Q, Benton MJ. 2012 The timing and pattern of biotic recovery following the 288 end-Permian mass extinction. *Nat. Geosc.* **5**, 375–383. 289 4 Bakker RT. 1977 Tetrapod mass extinctions – a model of the regulation of speciation 290 rates and immigration by cycles of topographic diversity. In *Patterns of Evolution as* 291 *Illustrated by the Fossil Record* (ed. Hallan A). New York, Elsevier. pp. 439–468. 292 5 Benton MJ, Tverdokhlebov VP, Surkov MV. 2004 Ecosystem remodelling among 293 vertebrates at the Permo-Triassic boundary in Russia. *Nature* **432**, 97–100. 294 6 Nesbitt SJ. 2011 The early evolution of archosaurs: relationships and the origin of 295 major clades. Bull. Am. Mus. Nat. Hist. 352, 1–292.

296	7	Brusatte SL, Nesbitt SJ, Irmis RB, Butler RJ, Benton MJ, Norell MA. 2010 The
297		origin and early radiation of dinosaurs. Earth Sci. Rev. 101, 68–100.
298	8	Langer MC, Ezcurra MD, Bittencourt J, Novas FE. 2010 The origin and early
299		evolution of dinosaurs. Biol. Rev. 85, 55–110.
300	9	Ezcurra MD, Scheyer TM, Butler RJ. 2014 The origin and early evolution of Sauria:
301		reassessing the Permian saurian fossil record and the timing of the crocodile-lizard
302		divergence. PLoS ONE 9, e89165.
303	10	Ezcurra MD. 2016 The phylogenetic relationships of basal archosauromorphs, with an
304		emphasis on the systematic of proterosuchian archosauriforms. <i>PeerJ</i> 4, e1778.
305	11	Brusatte SL, Benton MJ, Ruta M, Lloyd GT. 2008 Superiority, competition and
306		opportunism in the evolutionary radiation of dinosaurs. <i>Science</i> 321 , 1485–1488.
307	12	Butler RJ, Brusatte SL, Reich M, Nesbitt SJ, Schoch RR, Hornung JJ. 2011 The sail-
308		backed reptile Ctenosauriscus from the latest Early Triassic of Germany and the
309		timing and biogeography of the early archosaur radiation. <i>PLoS ONE</i> 6 , e25693.
310	13	Nesbitt SJ, Butler RJ, Ezcurra MD, Barrett PM, Stocker MR, Angielczyk KD, Smith
311		RMH, Sidor CA, Niedźwiedzki G, Sennikov A, Charig AJ. 2017 The earliest bird-line
312		archosaurs and the assembly of the dinosaur body plan. <i>Nature</i> 544 , 484–487.
313	14	Foth C, Ezcurra MD, Sookias R, Brusatte SL, Butler RJ. 2016 Unappreciated
314		diversification of stem archosaurs during the Middle Triassic predated the dominance
315		of dinosaurs. BMC Evol. Biol. 16, 188.
316	15	Norell, M. A. 1992 Taxic origin and temporal diversity: the effect of phylogeny. In
317		Extinction and Phylogeny (eds. Novacek MJ, Wheeler QD). New York, Columbia
318		University Press. pp. 88–118.



342	26	Bernardi M, Klein H, Petti FM, Ezcurra MD. 2015 The origin and early radiation of
343		archosauriforms: integrating the skeletal and footprint record. PLoS ONE 10,
344		e0128449.
345	27	Ezcurra MD, Butler RJ. 2015 Taxonomy of the proterosuchid archosauriforms
346		(Diapsida: Archosauromorpha) from the earliest Triassic of South Africa, and
347		implications for the early archosauriform radiation. <i>Palaeontology</i> 58 , 141–170.
348	28	Smith RMH, Botha-Brink J. 2014 Anatomy of a mass extinction: sedimentological
349		and taphonomic evidence for drought-induced die-offs at the Permo-Triassic
350		boundary in the main Karoo Basin, South Africa. Palaeogeogr. Palaeoclimatol.
351		Palaeoecol. 396 , 99–118.
352	29	Botha J, Smith RMH. 2006 Rapid vertebrate recuperation in the Karoo Basin of South
353		Africa following the end-Permian extinction. J. Afr. Earth Sci. 45, 502-514.
354	30	Irmis RB, Whiteside JH. 2011 Delayed recovery of non-marine tetrapods after the
355		end-Permian mass extinction tracks global carbon cycle. Proc. R. Soc. B 279,
356		1310-1318.
357	31	Button DJ, Lloyd GT, Ezcurra MD, Butler RJ. 2017 Mass extinctions drove global
358		faunal cosmopolitanism on the supercontinent Pangaea. Nat. Comm. 8, 733.
359	32	Payne JL, Lehrmann DJ, Wei J, Orchard MJ, Schrag DP, Knoll AH. 2004 Large
360		perturbations of the carbon cycle during recovery from the end-Permian extinction.
361		Science 305 , 506–509.
362	33	Whiteside JH, Ward PD. 2011 Ammonoid diversity and disparity track episodes of
363		chaotic carbon cycling during the early Mesozoic. <i>Geol.</i> 39 , 99–102.
364	34	Sun Y, Joachimski MM, Wignall PB, Yan C, Chen Y, Jiang H, Wang L, Lai X. 2012
365		Lethally hot temperatures during the Early Triassic greenhouse. Science 338,
366		366-370.



392 393 Figure captions 394 **Figure 1.** Diversity and abundance of late Permian—early Carnian archosauromorphs. (a) 395 Randomly selected, time-calibrated most parsimonious tree (MPT) showing the phylogenetic 396 diversity of early archosauromorphs. (b) Observed species count (red), phylogenetic diversity 397 (values from 10,000 randomly selected MPTs in grey; mean of those values in blue), and 398 number of individuals (green) per time bin. Silhouette labels in supplementary figure 11. 399 Avemet. = Avemetatarsalia. [two columns] 400 401 Figure 2. Evolutionary rates and morphological disparity of late Permian—early Carnian 402 archosauromorphs. (a) "Spaghetti" plot showing significantly fast (red) or slow (blue) rates 403 of phenotypic evolution calculated from 100 randomly selected, time-calibrated MPTs. Grey 404 points are non-significant values from the pooled average rate. Each thin line represents the 405 analysis of one MPT. Pie charts show the ratio of significantly fast (red), slow (blue), and 406 non-significant (white) rates at each time bin. (b) Morphological disparity of early 407 archosauromorphs represented by weighted mean pairwise dissimilarity (WMPD) generated 408 from GEDdm (green) and MORDdm (magenta), and its 95% confidence intervals generated 409 using 1,000 bootstrap replicates of the original data matrix. (c) Carbon isotope record from 410 the late Capitanian to the earliest Ladinian (taken from [30]). [two columns] 411 412 **Figure 3.** Morphospace occupation of late Permian–early Carnian archosauromorphs. (a-e)413 Sequence of morphospaces from the oldest to the youngest sampled time bin and (f) 414 morphospace of all time bins together. Each plot shows the first two principal coordinate 415 axes, which account for a summed variance of 18.23%. The black dots represent the position 416 in the morphospace of each terminal in that time bin and the grey dotted line represents the

convex hull of the morphospace of the previous time bin. The silhouettes show the
approximate position of different main clades in the morphospace (silhouette labels in
supplementary figure 11). Highly fragmentary taxa tend to occupy a position closer to $(x=0, 0)$
y=0) in the ordination of the GED dissimilarity matrix, and thus the high density of taxa in
this area is a methodological artefact (electronic supplementary material). [two columns]

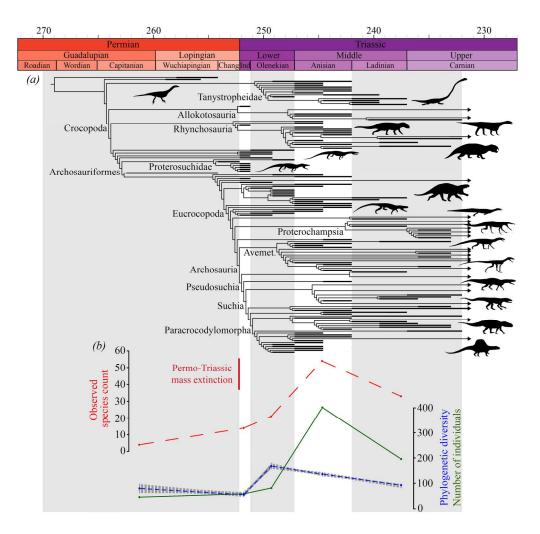


Figure 1. Diversity and abundance of late Permian—early Carnian archosauromorphs. (a) Randomly selected, time-calibrated most parsimonious tree (MPT) showing the phylogenetic diversity of early archosauromorphs. (b) Observed species count (red), phylogenetic diversity (values from 10,000 randomly selected MPTs in grey; mean of those values in blue), and number of individuals (green) per time bin. Silhouette labels in supplementary figure 11. Avemet. = Avemetatarsalia. [two columns]

175x171mm (300 x 300 DPI)

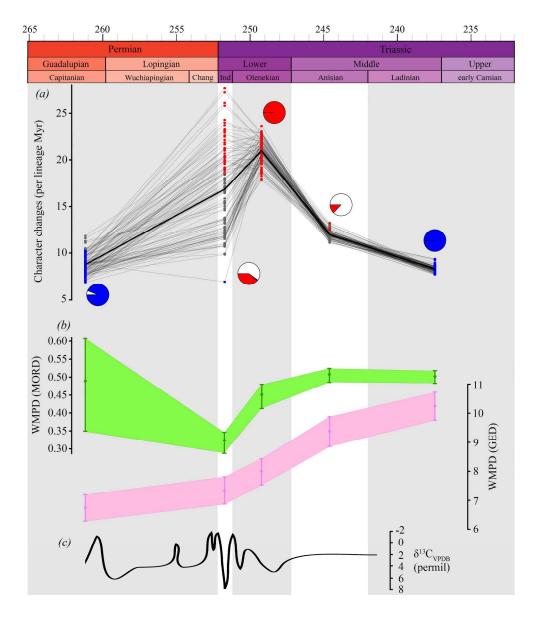


Figure 2. Evolutionary rates and morphological disparity of late Permian—early Carnian archosauromorphs. (a) "Spaghetti" plot showing significantly fast (red) or slow (blue) rates of phenotypic evolution calculated from 100 randomly selected, time-calibrated MPTs. Grey points are non-significant values from the pooled average rate. Each thin line represents the analysis of one MPT. Pie charts show the ratio of significantly fast (red), slow (blue), and non-significant (white) rates at each time bin. (b) Morphological disparity of early archosauromorphs represented by weighted mean pairwise dissimilarity (WMPD) generated from GEDdm (green) and MORDdm (magenta), and its 95% confidence intervals generated using 1,000 bootstrap replicates of the original data matrix. (c) Carbon isotope record from the late Capitanian to the earliest Ladinian (taken from [30]). [two columns]

200x228mm (300 x 300 DPI)

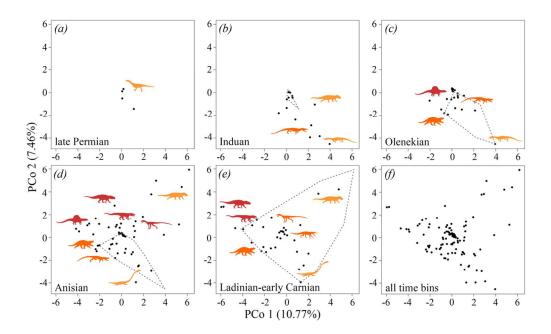


Figure 3. Morphospace occupation of late Permian—early Carnian archosauromorphs. (a—e) Sequence of morphospaces from the oldest to the youngest sampled time bin and (f) morphospace of all time bins together. Each plot shows the first two principal coordinate axes, which account for a summed variance of 18.23%. The black dots represent the position in the morphospace of each terminal in that time bin and the grey dotted line represents the convex hull of the morphospace of the previous time bin. The silhouettes show the approximate position of different main clades in the morphospace (silhouette labels in supplementary figure 11). Highly fragmentary taxa tend to occupy a position closer to (x=0, y=0) in the ordination of the GED dissimilarity matrix, and thus the high density of taxa in this area is a methodological artefact (electronic supplementary material). [two columns]

109x66mm (300 x 300 DPI)

- 1 **Table 1.** Results of the morphological disparity and evolutionary rates analyses. The disparity
- 2 metrics were calculated using GEDdm and MORDdm and their 95% confidence intervals
- were calculated based on 1,000 bootstrap replicates of the original data matrix. Reported
- 4 phylogenetic diversity and evolutionary rates are mean values and their respective standard
- 5 deviation. Evolutionary rate and weighted mean pairwise dissimilarity (WMPD) values that
- 6 significantly differ from those of the previous time bin are shown in bold. Car. = Carnian;
- 7 Evol. = Evolutionary; Lad. = Ladinian; ind. = individuals.

Time bin	Nº	Phylogenetic	WMPD (GED)	WMPD (MORD)	Evol. rates
	ind.	diversity			
late Permian	29	63.42±5.79	6.74(±6.26-7.21)	$0.489(\pm0.349-0.607)$	8.76±1.06
Induan	42	38.65±2.95	7.35(±6.89-7.79)	0.318(±0.288-0.346)	16.85±4.52
Olenekian	65	150.91±3.07	8.00(±7.54-8.43)	0.445(±0.412-0.480)	20.97±1.27
Anisian	383	119.52±2.10	9.38(±8.86-9.86)	0.505(±0.485-0.524)	11.97±0.46
Ladearly Car.	179	76.45±1.77	10.25(±9.70-10.74)	$0.501(\pm0.482-0.519)$	8.27±0.34