

New insights into landslide processes around volcanic islands from Remotely Operated Vehicle observations offshore Montserrat

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1 **New insights into landslide processes around volcanic islands from Remotely**
2 **Operated Vehicle (ROV) observations offshore Montserrat**

3

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19 **Abstract**

20 Submarine landslide deposits have been mapped around many volcanic islands, but
21 interpretations of their structure, composition and emplacement are hindered by the
22 challenges of investigating deposits directly. Here, we report on detailed observations of
23 four landslide deposits around Montserrat collected by Remotely Operated Vehicles,
24 integrating direct imagery and sampling with sediment-core and geophysical data. These
25 complementary approaches enable a more comprehensive view of large-scale mass
26 wasting processes around island-arc volcanoes than has been achievable previously. The
27 most recent landslide occurred at 11.5–14 ka (Deposit 1; 1.7 km³) and formed a radially-
28 spreading hummocky deposit that is morphologically similar to many subaerial debris-
29 avalanche deposits. Hummocks comprise angular lava and hydrothermally-altered
30 fragments, implying a deep-seated, central subaerial collapse, inferred to have removed a
31 major proportion of lavas from an eruptive period that now has little representation in the
32 subaerial volcanic record. A larger landslide (Deposit 2; 10 km³) occurred at ~130 ka and
33 transported intact fragments of the volcanic edifice, up to 900 m across and over 100 m
34 high. These fragments were rafted within the landslide, and are best exposed near the
35 margins of the deposit. The largest block preserves a primary stratigraphy of subaerial
36 volcanic breccias, of which the lower parts are encased in hemipelagic mud eroded from
37 the seafloor. Landslide deposits south of Montserrat (Deposits 3 and 5) indicate the wide
38 variety of debris-avalanche source lithologies around volcanic islands. Deposit 5
39 originated on the shallow submerged shelf, rather than the terrestrial volcanic edifice, and
40 is dominated by carbonate debris.

41

42 **1. Introduction**

43 Extensive submarine landslide deposits are common around volcanic islands [*Moore et*
44 *al.*, 1989; *Deplus et al.*, 2001; *Masson et al.*, 2002; *Coombs et al.*, 2007; *Silver et al.*,
45 2009]. Such landslides profoundly modify island morphology and affect the marine
46 environment through sudden deposition of material. They also pose major hazards
47 through direct inundation [*Siebert*, 1984], their potential association with explosive
48 volcanic blasts [*Bogoyavlenskaya et al.*, 1985], and tsunamis [*Ward and Day*, 2003;
49 *Satake*, 2007]. Much of our current understanding of large landslide deposits around
50 volcanic islands is based on geophysical surveys [e.g., *Deplus et al.*, 2001; *Coombs et al.*,
51 2007; *Watt et al.*, 2012a] and distal core samples of associated turbidites [*Hunt et al.*,
52 2011; *Trofimovs et al.*, 2013]. Only a few submarine volcanic landslide deposits have
53 been observed or sampled directly [*Yokose*, 2002; *Morgan et al.*, 2007; *Croff Bell et al.*,
54 2013; *Day et al.*, 2015]. Such observations provide structural and lithological information
55 relating to the landslide source and emplacement processes that cannot be obtained by
56 other means.

57 In this paper, we summarise results from two Remotely Operated Vehicle (ROV) surveys
58 of four landslide deposits offshore the volcanic island of Montserrat. Our aim is to
59 provide detailed information on the source (e.g., subaerial edifice, submarine flank,
60 surrounding seafloor), lithology (e.g., pyroclastic rock, dense lava, carbonate reef) and
61 structure (e.g., heterogeneous, disaggregated material; intact primary blocks) of material
62 within the deposits. This informs our understanding of the relationship between the
63 dominant lithology and morphology of landslide deposits [cf. *Masson et al.*, 2006] and
64 helps interpret landslide emplacement processes and interaction with the seafloor, which

65 is a significant control on the magnitude of landslide-generated tsunamis [Watt *et al.*,
66 2012a].

67

68 ***1.1. Data collection***

69 Two research expeditions of the RRS *James Cook* (JC83; March 2013) and the R/V
70 *Nautilus* (NA037; October 2013) deployed Remotely Operated Vehicles (ROVs) offshore
71 Montserrat to investigate submarine landslide deposits through high-definition video
72 filming, still images, and a remotely manipulated sampling arm. Expedition JC83
73 deployed the Isis ROV, collecting footage during four dives SE of Montserrat (Figure 1;
74 Isis dive numbers are prefixed I). Dimensions of outcrops and rocks were estimated using
75 two laser points in the ROV field of view, which are 10 cm apart. A vibrocore attachment
76 collected a single core during Dive I213, but this attachment, as well as the manipulator
77 arm, was not operational during the remainder of the cruise. Expedition NA037 [Carey *et al.*
78 *et al.*, 2014] deployed a two-vehicle ROV system (Hercules and Argus) during three dives
79 south and east of Montserrat. In addition to imagery, it collected 61 samples via a
80 manipulator arm (Figure 1; Hercules/Argus dive numbers are prefixed H). The largest
81 rocks or consolidated-sediment samples that could be collected were 20 cm in diameter.
82 ROV-based technology has been used in Hawaii to investigate submarine volcanic-island
83 landslide processes [Yokose, 2002; Coombs *et al.*, 2004; Yokose and Lipman, 2004;
84 Morgan *et al.*, 2007], but our work is among the first to apply such methods elsewhere
85 [cf. Croff Bell *et al.*, 2013].

86

87 ***1.2. Terminology***

88 Following past studies around volcanic islands [e.g., *Moore et al.*, 1989; *Masson et al.*,
89 2002] we use *landslide* as a general term for any slope failure and the resulting mass
90 movement. The landslide deposits described here originated as failures of rock on the
91 subaerial and submerged island flanks, which fragmented to form a *debris avalanche*,
92 where the disintegrating mass is dispersed between clearly defined source and
93 depositional regions. Progressive fragmentation and spreading results in the characteristic
94 hummocky topography of debris-avalanche deposits [*Siebert*, 1984; *Glicken*, 1996;
95 *Paguican et al.*, 2014], but the specific character of the debris avalanche (and its deposit)
96 may depend on the nature of material within the landslide (e.g., density, strength,
97 homogeneity) [*Naranjo and Francis*, 1987; *Masson et al.*, 2006; *Dufresne and Davies*,
98 2009; *Watt et al.*, 2014]. Debris avalanches originating in clay-rich terrains, such as
99 hydrothermally altered portions of volcanic edifices, may be relatively cohesive. The
100 incorporation of basal sediment (e.g., hemipelagic mud from the seafloor) may also
101 promote more cohesive flow characteristics. For simplicity, we use *debris-avalanche*
102 *deposit* to refer to all deposits, rich in volcanic rock fragments, that directly result from
103 the initial landslide. In marine environments, seafloor-sediment failure [*Watt et al.*,
104 2012b, 2014] associated with debris-avalanche emplacement may produce more
105 extensive deposits. In addition, landslides around volcanic islands may generate dilute
106 and highly mobile *turbidity currents* [*Talling et al.*, 2012] from the mixing of primary
107 landslide material or disrupted marine sediment with seawater, depositing *turbidites*.

108

109 **2. Study region**

110 Montserrat is located in the northern Lesser Antilles Arc and comprises four volcanic
111 centers dating back to at least 2.5 Ma (Figure 1) [Harford *et al.*, 2002]. The andesitic
112 Soufrière Hills volcano has been active since 250 ka [Harford *et al.*, 2002, Smith *et al.*,
113 2007], interrupted by a short episode of basaltic volcanism at ~130 ka that formed the
114 South Soufrière Hills center. An important aspect of the geological history of Soufrière
115 Hills (and of Montserrat in general) is the occurrence of large landslides. Several debris-
116 avalanche deposits, with volumes between 0.3 and 10 km³, have been identified offshore
117 southern Montserrat from geophysical surveys [Le Friant *et al.*, 2004; Lebas *et al.*, 2011;
118 Watt *et al.*, 2012a,b]. In addition to these surveys, the identification and correlation of
119 tephra fall deposits and turbidites within marine sediment cores provides a detailed record
120 of past activity on the island [Le Friant *et al.*, 2009, 2015; Trofimovs *et al.*, 2013; Cassidy
121 *et al.*, 2013; Wall-Palmer *et al.*, 2014]. These studies provide age constraints on landslide
122 deposits and contribute to understanding the context of major landslides in the broader
123 volcanic history of the island. However, direct core sampling of the block-rich volcanic
124 landslide deposits has been unsuccessful, because of their coarse and heterogeneous
125 nature.

126 The 1995-to-recent eruption of Soufrière Hills has involved the growth and collapse of a
127 series of andesitic lava domes, generating pyroclastic flows [Wadge *et al.*, 2014]. The
128 largest dome collapse, in 2003, involved >0.21 km³ of material [Herd *et al.*, 2005]. East
129 of Montserrat, submarine deposits from several collapse-driven pyroclastic flows have
130 formed lobes with a cumulative thickness of 100 m, extending 7 km from the coastline
131 [Figure 1; Trofimovs *et al.*, 2008; Le Friant *et al.*, 2009].

132

133 **2.1. Terrestrial morphology and landslide scars**

134 Prior to its recent activity, Soufrière Hills consisted of a series of lava domes surrounding
135 a prominent crescent-shaped collapse scar (English's Crater). This scar was open to the
136 east and led directly into the Tar River valley (Figure 1). English's Crater has been the
137 location of lava extrusion since 1995, and is presently occupied by a lava dome with a
138 volume of $>0.19 \text{ km}^3$ [Stinton *et al.*, 2014]. Dating of material within English's Crater
139 shows that two eruptive or mass-wasting events, of unconstrained size, occurred at $\sim 2 \text{ ka}$
140 and $\sim 6 \text{ ka}$ [Smith *et al.*, 2007; Boudon *et al.*, 2007]. This indicates that the crater formed
141 at $\geq 6 \text{ ka}$.

142 East of the Tar River valley, a 3.5-km-wide chute is cut into the submerged SE flank of
143 Montserrat [Figure 1; Le Friant *et al.*, 2004]. This chute is attributed to a large landslide
144 that formed an elongate offshore deposit named Deposit 2 [Le Friant *et al.*, 2004]. Within
145 the northern part of the chute, a 1.2-km-wide depression aligns closely with the Tar River
146 valley and English's Crater. Collectively, these structures may mark the source and
147 pathway of an offshore landslide deposit named Deposit 1 [Le Friant *et al.*, 2004; Lebas
148 *et al.*, 2011]. Deposit 1 has a volume of 1.7 km^3 , whilst English's Crater represents ~ 0.5
149 km^3 of missing rock [Le Friant *et al.*, 2004]. The submerged chute has a volume of ~ 0.5
150 to 1.1 km^3 [Watt *et al.*, 2012b] but may be partly infilled by later aggradation.
151 Notwithstanding the large uncertainties (owing, for example, to a lack of constraints on
152 pre-existing topography), these estimated volumes suggest that Deposit 1 comprises both
153 subaerial material from English's Crater and submerged material from the northern part
154 of the chute. A reduced bulk density and seafloor-sediment incorporation may account for
155 some increase in the deposit volume versus the inferred failure volume.

156 Two further landslide deposits, termed Deposits 3 and 5, are located south of Montserrat
157 (Figure 1; note that Deposit 4 is buried beneath Deposit 3 and is not discussed further
158 here). These deposits align with scars in the island shelf but are not associated with any
159 visible subaerial collapse structures.

160

161 ***2.2. Morphological description of landslide deposits***

162 Deposits 1, 2, 3 and 5 are all defined by mounded, irregular areas of seafloor (Figure 1).
163 Within each deposit, the mounded surface may either represent hummocks – hills of
164 amalgamated landslide material, typical of subaerial debris-avalanche deposits [Siebert,
165 1984] – or individual scattered blocks, representing largely intact fragments of the initial
166 landslide mass [cf. Watt *et al.*, 2014].

167

168 ***2.2.1. Deposit 1***

169 The margin of Deposit 1 is defined as the limit of a hummocky, fan-shaped deposit that
170 extends 10.5 km offshore the Tar River valley, to water depths of 1000 m, and covers ~50
171 km². The deposit contains many tens of hummocks that are up to 200 m long and
172 protrude tens of meters above surrounding seafloor. The hummocks are evenly
173 distributed, without preferential accumulation at the margins or center of the deposit.
174 Seismic reflection data resolve no prominent internal structures within Deposit 1
175 [Crutchley *et al.*, 2013; Karstens *et al.*, 2013].

176

177 ***2.2.2 Deposit 2***

178 Deposit 2 is partially buried beneath Deposit 1 and is more extensive and voluminous
179 than the other deposits considered here, comprising $\sim 10 \text{ km}^3$ of material [Lebas *et al.*,
180 2011; Watt *et al.*, 2012a,b]. It has been proposed that the central, blocky part of Deposit 2
181 originated as a collapse of the volcanic edifice, which then triggered extensive failure of
182 the surrounding seafloor sediment [Watt *et al.*, 2012b; 2014]. IODP drilling (Figure 1)
183 confirms that the distal part of Deposit 2 comprises seafloor sediment [Le Friant *et al.*,
184 2015].

185 Here, we attribute the notably large blocks to the east of Montserrat to Deposit 2 (Figure
186 1), based on interpretations of available seismic and bathymetric data [Watt *et al.*, 2012b].
187 The most prominent of these blocks lies close to the eastern margin of Deposit 1, and has
188 an angular, steep-sided form that contrasts with the rounded hummocks of Deposit 1. It is
189 900 m long, 700 m wide and 100 m high, and may have a similar buried extent,
190 indicating a total volume of $\sim 0.05\text{--}0.08 \text{ km}^3$ [Crutchley *et al.*, 2013]. To place this
191 volume into context, it is approximately ten times that of Wembley Stadium in London
192 (0.004 km^3), one of the world's largest sports grounds. A 2-km arc of blocks with
193 comparable dimensions to the "Wembley" block (as it is referred to here) marks the
194 proximal southern margin of Deposit 2 (Figure 1). More very large blocks or hummocks
195 occur further east, within the central part of Deposit 2, but are partially buried by younger
196 sediment.

197

198 2.2.3. Deposit 3

199 Deposit 3 extends 10.5 km to the south of Montserrat, reaching water depths of 950 m.
200 Seismic reflection profiles suggest that it is thinner than Deposit 1, and mainly comprises

201 scattered large blocks [*Lebas et al.*, 2011; *Watt et al.*, 2012b] with a total volume of <1
202 km³.

203

204 *2.2.4. Deposit 5*

205 Deposit 5 has a poorly constrained volume of ~0.3 km³ [*Le Friant et al.*, 2004] and is
206 associated with a scar on the submerged coastal shelf on the south-western side of
207 Montserrat. It is defined by a hummocky field of debris that can be traced 7 km offshore
208 to a water depth of about 830 m.

209

210 *2.3. Ages of landslide deposits*

211 Dating of submarine landslide deposits is best achieved by constraining the age and
212 accumulation rate of hemipelagic sediment both above and below the deposit. However,
213 given the difficulties of coring through landslide deposits, ages are often based either on
214 the oldest sediment overlying the deposits or on the age of turbidites that have been
215 correlated with them. In the former approach, the distance between the base of a sediment
216 core and the top of the landslide deposit may be unknown, and any age thus derived is a
217 minimum. In the latter approach, it is potentially difficult to correlate a specific turbidite
218 with a landslide deposit, given that neither necessarily has a unique composition in terms
219 of chemistry or componentry.

220

221 *2.3.1. Deposit 1*

222 The best direct age constraint for Deposit 1 comes from core JR123-54 (collected in
223 2005; Figures 1 and 2) [*Trofimovs et al.*, 2013], located on a hummock. The basal unit in

224 the core is a mixed bioclastic and volcanoclastic turbidite, the lowest part of which
225 comprises poorly-sorted gravel containing altered lava clasts, which may correspond to
226 the top surface of Deposit 1 [Trofimovs *et al.*, 2013]. Multiple radiocarbon dates (Table 1)
227 indicate an age of ~11.5 ka for this turbidite (a potentially bioturbated sample within the
228 uppermost part of the turbidite provides a maximum age of 12.3 ka).

229 Deposit 1 may correlate with a large ($>0.4 \text{ km}^3$) turbidite that extends over 30 km to the
230 south of Montserrat (Figure 1), dated by multiple radiocarbon ages at 12–14 ka
231 [Trofimovs *et al.*, 2013]. The turbidite is by far the largest-volume and most erosive event
232 in the offshore stratigraphy during the past 110 ka, and its thickest part coincides with the
233 margin of Deposit 1. The timing, distribution and magnitude of the two deposits thus
234 support their correlation. The stratigraphy of the turbidite is complex and spatially
235 variable [Trofimovs *et al.*, 2010], but taken as a whole it comprises equal proportions of
236 biological (calcium carbonate) and volcanic clasts. This contrasts with turbidites derived
237 from pyroclastic flows in the present eruption of Soufrière Hills, which are >95%
238 volcanoclastic [Trofimovs *et al.*, 2008]. Thus, the source event of the 12–14 ka turbidite
239 must have mobilized a significant proportion of submarine, carbonate-rich material,
240 either by contemporaneous failure and disaggregation of carbonate-rich lithologies (i.e.
241 from the island's carbonate shelf), or by erosion of carbonate-rich seafloor sediment.

242 Combining the age determinations from JR123-54 and the mixed turbidite, Deposit 1
243 occurred at 11.5–14 ka.

244

245 *2.3.2. Deposit 2*

246 Sediment cores from IODP Expedition 340 (Figure 1) [*Le Friant et al.*, 2015] place the
247 top of Deposit 2 at ~130 ka [*Cassidy et al.*, 2015], based both on oxygen isotope
248 stratigraphy of younger hemipelagic mud and on the correlation of basaltic deposits,
249 which immediately overlie Deposit 2, with volcanism at South Soufrière Hills (dated at
250 130 ka by Ar-Ar ages of subaerial lavas [*Harford et al.*, 2002]). This age is consistent
251 with an earlier estimate of ~140 ka derived from regional sediment accumulation rates
252 [*Watt et al.*, 2012b].

253

254 *2.3.3. Deposit 3*

255 A spatial correlation with a mafic volcanoclastic turbidite [*Cassidy et al.*, 2014], dated at
256 60–130 ka, provides a possible age constraint for Deposit 3. If correct, the correlation
257 implies a mafic source lithology for the landslide. Seismic reflection profiles indicate a
258 sedimentary cover of 5–10 m over Deposit 3, implying an age of 100–200 ka [based on
259 local sedimentation rates of 0.05 m kyr^{-1} ; *Watt et al.*, 2012b].

260

261 *2.3.4. Deposit 5*

262 The thickest part of a mixed volcanoclastic and bioclastic turbidite is co-located with
263 Deposit 5, suggesting a correlation between the two deposits [*Cassidy et al.*, 2013]. The
264 high bioclastic content of the turbidite is consistent with the identified landslide source
265 scar on the submerged coastal shelf. The turbidite has an erosive base in hemipelagic
266 sediment dated at 35 ka, and lies directly beneath a volcanoclastic turbidite dated at 8–12
267 ka. Deposit 5 is therefore similar in age to Deposit 1. The cluster of landslide and

268 turbidite deposits at 8–14 ka suggests a period of relatively heightened mass-wasting
269 activity at Montserrat.

270

271 **3. New ROV-based Observations**

272 The principal ROV observations for each landslide deposit are described and interpreted
273 in this section. This interpretation draws on data from pre-existing core samples and
274 geophysical data. More specific discussion of landslide processes relating to Deposits 1
275 and 2 is provided in Section 4. In addition to the figures described here, short video files
276 of key exposures are provided as Supporting Information.

277

278 ***3.1. Deposit 1***

279 *3.1.1. Hummock exposures*

280 ROV observations made on seven hummocks in Deposit 1 (Figure 1) indicate broadly
281 similar mixtures of lithologies, with representative images shown in Figure 2. The top of
282 individual hummocks provide the best outcrops; a talus of scattered rocks and partially
283 eroded sedimentary drape obscure surrounding slopes. Outcrops expose volcanic breccia,
284 with wide variation in grain size, sorting, presence or absence of a fine matrix, presence
285 or absence of layering, clast shape and alteration. Lithologically diverse domains occur at
286 a range of scales, both within and between hummocks.

287 A poorly sorted and matrix-supported breccia is the dominant lithology, displaying a
288 range of colorations and with generally sharp, but occasionally diffuse, irregular
289 boundaries between colored domains. Pale colored domains are interpreted as
290 hydrothermally altered volcanic breccias; the diverse coloration (white and pale-yellow

291 are the most common, but green, yellow, orange and brown also occur) indicates a range
292 of mineral assemblages, and suggests that different zones of hypogene alteration in the
293 failure region [cf. *John et al.*, 2008] were efficiently mixed during debris-avalanche
294 emplacement. Undulose boundaries (Figure 2c,d) indicate shearing and stretching of
295 altered domains during transport.

296 Altered breccias often lie in direct contact with dark gray, monomict, clast-supported to
297 marginally matrix-supported breccias. Clasts are angular to sub-angular and vary in size
298 from a few meters to a centimeter (Figure 2b). This lithology is interpreted as unaltered
299 autoclastic breccia associated with lava dome extrusion. Pink to red lava breccias also
300 occur, with otherwise similar characteristics to the monomict gray breccias, and are
301 indicative of hematite formed in a subaerial setting. In one case (Figure 2e), narrow (10–
302 30 cm) and irregular zones of alteration were observed passing through a large outcrop of
303 gray lava breccias.

304 Samples of the dense lavas (NA037-008 and -011; see Supporting Information) show a
305 phenocryst assemblage dominated by plagioclase and orthopyroxene, with frequent
306 amphibole largely replaced by an alteration assemblage. This assemblage is typical of
307 Soufrière Hills andesites erupted since ~110 ka [*Harford et al.*, 2002]. We identified no
308 unequivocal biological (carbonate) material or structures within Deposit 1. A sample of
309 orange-brown hydrothermally altered rock (NA037-009; Figure 2a) contained abundant
310 clay minerals and hydrothermally altered ferromagnesian and feldspar crystals.

311

312 *3.1.2. Deposit 1 sedimentary drape*

313 The sedimentary drape that overlies Deposit 1 is well exposed on the sides of several
314 hummocks, where it has been eroded by bottom currents or local slope failures (Figure
315 3). Interpretations of these exposures have drawn on the extensive previous core
316 sampling of the top ~5 m of seafloor sediment in the area, which comprises an
317 interbedded sequence of hemipelagic mud and volcanoclastic, bioclastic or mixed
318 turbidites [JR123; *Trofimovs et al.*, 2010, 2013].

319 The observed exposures comprise a mixture of fine-grained, white to pale-gray
320 hemipelagic sediment and interbedded sandy turbidites. Hemipelagic mud intervals
321 frequently contain coarse volcanic clasts (Figure 3), which are likely to be locally derived
322 (e.g. by reworking from upslope on a hummock). These poorly-sorted beds of outsized
323 volcanic clasts set in hemipelagic mud are similar to the talus deposits at the base of the
324 SW Wembley-block exposures (Section 3.2.1; Figure 4d). Bed dips are parallel to the
325 local slope, and sometimes up to 40° (Figure 3b). These heterogeneous beds were not
326 sampled by the JR123 cores, but we note that some attempts at coring failed, perhaps due
327 to the coarse nature of this material.

328 In several exposures, the basal unit of the drape (i.e. the deposit immediately overlying
329 Deposit 1) is a well-sorted, monomict and clast-supported, matrix-free volcanic breccia
330 of dense, gray cm-scale andesite clasts. This unit appears to be relatively continuous over
331 Deposit 1 (Figure 3e). This immature, matrix-free breccia is similar to beds found within
332 volcanic blast deposits on the surface of some subaerial debris-avalanche deposits
333 [*Hoblitt et al.*, 1981; *Bogoyavlenskaya et al.*, 1985; *Clavero et al.*, 2004; *Belousov et al.*,
334 2007], and provides possible evidence of a lateral explosion accompanying the Deposit 1
335 landslide. An alternative possibility is that this unit represents a capping, coarse-grained

336 turbidite generated by the debris avalanche; it may correlate with the gray volcanoclastic
337 beds in the widespread 12–14 ka turbidite [cf. *Trofimovs et al.*, 2013].

338

339 **3.2. Deposit 2**

340 *3.2.1. Wembley block*

341 The Wembley block differs from the hummocks within Deposit 1 in its scale,
342 componentry and shape. It also displays some differences in post-emplacment
343 sedimentary cover. Its angular, steep-sided form suggests that it is a single fragment of
344 the volcanic edifice. The exposed base of the block is not its true base, which may be as
345 much as 100 m below the seafloor [cf. *Crutchley et al.*, 2013].

346

347 *3.2.1.1. Surface exposures*

348 Continuous exposures on the SE side of the Wembley block are summarized in Figure 5
349 (Dive I217). The lower half of the block exposes a largely structureless breccia of
350 angular, dense, gray andesite clasts set within a uniform, white to pale-gray fine-grained
351 matrix, which erodes with a sculpted, pitted appearance (Figure 4a,b,f). We interpret this
352 matrix as hemipelagic mud, because of its similar appearance to the hemipelagite
353 exposed in scarps that cut the seafloor east of the block (this mud has been sampled in
354 numerous cores [*Trofimovs et al.*, 2013]). The exposures change abruptly 26 m above the
355 seafloor, to volcanic breccias of dense angular clasts, either gray or red in color,
356 displaying crude low-angle bedding (Figure 6d), but without any pale mud matrix (Figure
357 5). The volcanic breccias are similar in appearance to unaltered breccias in Deposit 1, but
358 hydrothermally altered rocks are absent. Some clasts show fractures (Figure 6e) that may

359 reflect in-situ brecciation acquired by vibration and collision during transport. Exposures
360 vary from matrix- to clast-supported breccias. Although most are monomict, some beds
361 contain mixtures of gray and red lava fragments, and are sub-rounded in parts. We
362 interpret the monomict breccias as dome-collapse block-and-ash flow deposits, and the
363 more mixed, rounded units, as reworking of the same material. The common occurrence
364 of reddened lavas suggests a subaerial origin.

365 Very dark lava clasts are exposed near the base of the ESE side of the Wembley block
366 (Dive I217). Based on samples with a similar appearance from Deposit 3, we interpret
367 these as blocks with ferromanganese surface encrustation (Figure 4e, 6c). Such
368 encrustation is likely to have formed after deposition, assuming that the block surfaces
369 were not previously exposed in a submarine environment. It is unclear why this
370 encrustation is restricted to a single part of the Wembley block, but the formation of
371 ferromanganese crusts can be strongly dependent on water depth and local biological
372 activity [*Hodkinson and Cronan, 1991*].

373 The base of the Wembley block on its SW side (Dive H1308) also exposes volcanic
374 breccias within a hemipelagic mud matrix, but here they display crude, high angle
375 bedding, and unconformably overlie a monomict volcanic breccia without any mud
376 matrix (Figure 4d). We interpret the bedded mud-supported breccia as a post-
377 emplacement talus of volcanic clasts mixed with continuously depositing hemipelagic
378 sediment, derived from periodic mass wasting of the steep slopes of the Wembley block.
379 The monomict breccia is thus the surface of the primary block. Higher up the SW side of
380 the block, clast-supported volcanic breccias dominate (Figure 6a). Overall, these are more
381 angular than the breccias on the SE side. We interpret the whole sequence as autoclastic

382 and reworked lava breccias forming as talus around an active lava dome. The greater
383 prevalence of reworked breccias on the SE side of the block suggests a more marginal
384 facies than those on the SW, which is plausible given the 900-m dimensions of the block.
385 The entire block is thus a fragment of the subaerial volcano, transported intact to its
386 present position.

387

388 3.2.1.2. Seafloor interaction

389 Although the mud-supported breccias on the SW side of the block are clearly post-
390 emplacement talus deposits, the mud-supported breccias on the SE side may be a syn-
391 emplacement feature. Here, the mud matrix is present on sub-vertical and highly
392 irregular, gullied slopes, sometimes showing a gradational contact with monomict, clast-
393 supported volcanic breccias (Figure 4), and is prevalent below a sharp and broadly
394 horizontal boundary. The SE side of the block was the frontal section during block
395 emplacement, and seismic reflection data indicate that the emplacement of Deposit 2
396 involved substantial erosion of seafloor sediment [Watt *et al.*, 2012a,b]. Incorporation of
397 mud into the brecciated surface of the block may have occurred during this process,
398 explaining the presence of this matrix in the lower and frontal part of the block. This
399 sediment injection is not necessarily deeply penetrating. We favour this interpretation
400 over alternative origins for the marine sediment matrix on the SE side of the Wembley
401 block. Hemipelagic mud characterizes marine sedimentation on the deep seafloor around
402 Montserrat; if a marine matrix was a primary characteristic of the block (and if we
403 assume the block originated on the submerged island flanks), we would expect more
404 evidence of shallow water carbonate rocks, and for the volcanic breccias to be more

405 extensively reworked. Rare white fragments are observed in the hemipelagic mud (Figure
406 5), up to 2 cm across, but these may be deep water bivalves of the type observed (up to
407 0.5 cm across) on the south side of Montserrat.

408

409 3.2.1.3. Sample descriptions

410 A single lava sample from the block (NA037-001; see Supporting Information)
411 comprises fresh, dense porphyritic andesite with a phenocryst assemblage of plagioclase,
412 orthopyroxene and clinopyroxene. Hornblende is absent. This assemblage contrasts with
413 the andesite mineralogy that has predominated on Montserrat since ~110 ka (and that
414 occurs in Deposit 1), but is similar to rocks erupted before 130 ka [*Harford et al.*, 2002;
415 *Zellmer et al.*, 2003].

416 Loose yellow clasts of highly indurated carbonate, up to 30 cm across, were observed on
417 the block surface near the top of the SW side of the block (Dive H1308; Figure 7). A
418 sample of this material (NA037-002; see Supporting Information) is a coralgal limestone
419 consisting of a mixture of large (cm-sized) rhodoliths, benthic foraminifera (notably
420 *Amphistegina* and peneroplids) and other bioclasts (including gastropods, bivalves,
421 echinoids and calcareous red algal fragments) within a matrix of micrite. Microbialite-
422 micritic filaments and peloids probably represent in-situ bacterial precipitates. Some
423 bioclasts have textures indicating replacement of original aragonite by neomorphic
424 calcite. The characteristics of this clast suggest formation at shelfal depths, but the
425 replacement of aragonite suggests diagenesis either in a meteoric environment or in its
426 current deep-water setting (900 m). A second sample (NA037-005) is a weakly indurated
427 micritic limestone with planktonic foraminifera (*Globorotalia*, *Orbulina*), planktonic

428 gastropods (including pteropods), minor fragments of shallow-water bioclasts (bivalves,
429 foraminifera, echinoids), and silt-sized volcanic crystals set in a micrite matrix with
430 conspicuous (mm-sized) burrow fills. The sample exterior has some tubeworm clasts and
431 small coral fragments. The mix of shallow and deep water fauna, with incorporation of
432 minor volcanic fragments and aragonite replacement all suggest transport from a shallow
433 to a deeper environment. We infer that these clasts were transported from shallow water
434 to their current position during emplacement of the Wembley block. They may represent
435 material from the submarine shelf that was eroded during the passage of the volcanic
436 debris avalanche, which fell onto the surface of the block before being transported to
437 their present position.

438

439 *3.2.2. Large southern block*

440 A large block south of Deposit 1, mapped as a marginal block within Deposit 2 [Figure 1,
441 Dive I213; *Watt et al.*, 2012b], comprises monomict lava breccias with dark coloration,
442 interpreted as ferromanganese encrustation. Gray volcanoclastic sand from the recent
443 Soufrière Hills eruption obscures much of the block surface. Our limited observations
444 suggest that the block is lithologically similar to the Wembley block.

445

446 *3.2.3. Wembley block sedimentary drape*

447 Approximately 3 m of marine sediment is exposed on top of the SE side of the Wembley
448 block (Figure 5). Prominent beds of white hemipelagic mud are interbedded with three
449 thicker, recessive gray sandy units, interpreted as turbidites, which are partly obscured by
450 deposits of recent volcanoclastic sand (Figure 6f). In comparison with the stratigraphy of

451 core JR123-21, collected on top of the Wembley block in 2005 [Trofimovs *et al.*, 2008,
452 2010], the drape on the SE edge of the block contains thicker turbidites and thinner
453 hemipelagite intervals (Figure 8). Both sequences are very different in terms of both layer
454 thickness and characteristics from the stratigraphy recovered in over 20 vibracores from
455 the surrounding seafloor [JR123; Trofimovs *et al.*, 2008, 2010, 2013] (Figure 8).

456 The youngest turbidites in the correlated stratigraphy from the surrounding seafloor are
457 much thicker than those from JR123-21. This may be explained by the elevated position
458 of the block, where clast concentration in turbidity currents may have been lower
459 (resulting in thinner deposits). However, the sandy beds at the base of JR123-21 are
460 notably thick. These lower units are almost purely volcanoclastic, and do not correlate
461 clearly with any turbidites in the local stratigraphy, which is well defined at ages <110 ka
462 [Trofimovs *et al.*, 2013]. They may be the deposits of older turbidity currents generated
463 during the emplacement of Deposit 2.

464 The Wembley block is mapped as part of Deposit 2 [Watt *et al.*, 2012a,b; Crutchley *et al.*,
465 2013], but its location (Figure 1) suggests that it could be an outrunner block within
466 Deposit 1. Seismic reflection profiles and the regional turbidite record provide no
467 evidence of major landslides in the period between Deposits 2 (~130 ka) and Deposit 1
468 (11.5–14 ka). New radiocarbon dates from JR123-21 (Figure 8; Table 1) extend beyond
469 the limits of radiocarbon dating (43.5 ka), supporting interpretation of the Wembley block
470 as part of Deposit 2. However, the dates do not provide good constraints on turbidite ages
471 or hemipelagic sedimentation rates, because several ages cluster around 43 ka, and some
472 are out of stratigraphic sequence (Figure 8). This suggests extensive bioturbation or the
473 possible reworking of material derived from bioclastic turbidites with background

474 hemipelagic sediment. The 1.2-m thickness of hemipelagic intervals in JR123-21 also
475 supports a pre-Deposit 1 age for the Wembley block: post-Deposit 1 hemipelagic mud on
476 the surrounding seafloor has a cumulative thickness of 70–80 cm; and hemipelagic
477 sedimentation rates of 6.6 cm kyr^{-1} , estimated from a 45-cm vibrocore (JC83-VC1) on top
478 of the large southern block (Figure 1; Table 1) imply that the hemipelagite in JR123-21
479 represents >18 kyr. However, the sedimentary drape is surprisingly thin if the
480 emplacement age of the block is 130 ka. Thus, although the balance of observations
481 suggests that the Wembley block lies within Deposit 2, several aspects of the sedimentary
482 drape remain puzzling.

483

484 **3.3. Deposit 3**

485 The surface of Deposit 3 (Dive H1310; Figure 1) is not well exposed, but occasional
486 clusters of meter-scale blocks, with features such as well-developed radial jointing
487 (Figure 9a), protrude through younger sedimentary cover. The blocks are dense
488 porphyritic andesite lavas with a very dark surface coating, caused by thick (up to 3 mm)
489 manganese encrustations. Examination of two thin sections (NA037-037 and -042;
490 Supporting Information) indicates a phenocryst assemblage of plagioclase, clinopyroxene
491 and orthopyroxene. Orthopyroxene is less abundant than in the Wembley block sample
492 (NA037-001). The assemblage is comparable to that observed in the pre-130 ka andesites
493 of Soufrière Hills and in some of South Soufrière Hills rocks [Zellmer *et al.*, 2003],
494 although olivine is absent. An origin from South Soufrière Hills would be consistent the
495 previous correlation of Deposit 3 with a mafic volcanoclastic turbidite [Cassidy *et al.*,
496 2014]. The prevalence of angular, fractured lava blocks suggests a subaerial source for

497 the landslide; the absence of a visible source scar and a lack hydrothermally altered
498 material in the exposures suggests that this landslide may have been relatively shallow-
499 seated.

500

501 **3.4. Deposit 5**

502 Clusters of blocks in Deposit 5 are well exposed at depths of 750–830 m (Dive H1309;
503 Figure 1). Blocks comprise massive carbonate fragments (Figure 9f) and well-bedded
504 carbonate-cemented volcanoclastic conglomerates. The well-rounded conglomerates
505 (Figure 9c) are comparable to beach cobbles and mature fluvial deposits, and the
506 carbonate fragments are similar to large slabs of hardground observed in separate dives at
507 depths of 100–200 meters off the southern coast of Montserrat. A single large slab of reef
508 rock has karstic features (deeply incised channels) indicative of subaerial exposure,
509 perhaps during a low stand in sea level (Figure 9d,e).

510 One carbonate sample (NA037-026; Figure 10, Supporting Information) is a dense
511 limestone of encrusted volcanic clasts and bioclasts, including benthic and planktonic
512 foraminifera, calcareous red algae, mollusc fragments, serpulids, sponge spicules,
513 radiolaria, echinoid spines and pteropods, cemented by micritic-microsparitic-sparry
514 calcite cement. The encrusted grains (comparable to oncoids or rhodoliths) probably
515 formed by rolling in intermittent currents in shallow to moderate water environments,
516 consistent with the fossil assemblage. Encrusting foraminifera on red algal crust occur
517 with microbial filaments. Aragonitic gastropod and sponge fragments are replaced by
518 coarse calcite, consistent with diagenetic alteration following transport to a deep-water
519 environment. Phosphate grains of probable microbial origin occur within cavities (sponge

520 borings) in calcareous algae. A further sample (NA037-025) is a well-sorted, porous
521 cemented bioclastic grainstone (medium to coarse sand) cemented by thin (20 to 50 μm)
522 isopachous bladed calcite. Grains include shallow-water foraminifera (penerolids),
523 calcareous algae (branched forms), green algae (*Halimeda*), minor bivalve fragments and
524 volcanic clasts. Areas of peloidal sediment are likely to be the result of bacterial
525 precipitation. Our observations support the previous conclusion [*Le Friant et al.*, 2004;
526 *Cassidy et al.*, 2013] that Deposit 5 originated as a shallow-seated collapse of the coastal
527 shelf.

528

529 ***3.5. Sharp-faced depressions in young sediment***

530 Numerous sharp-faced depressions, up to a few meters deep, occur on the seafloor
531 between hummocks in Deposit 5 and to the east of Deposit 1 [cf. *Watt et al.*, 2012b].
532 These structures are defined by arcuate scarps, in some cases forming fully enclosed,
533 round depressions, exposing near-vertical cliffs through the seafloor sedimentary
534 sequence (Figure 11b,c). The depressions are at least tens of meters across in the vicinity
535 of Deposit 5, and up to hundreds of meters across to the east of Deposit 1. The
536 stratigraphy of scarps east of Deposit 1 (Figure 11c) comprises interbedded turbidites and
537 hemipelagic mud but is difficult to correlate precisely with the regional turbidite
538 stratigraphy (Figure 8). The good exposure of the scarps suggests that they cut through to
539 the youngest Holocene deposits and that they therefore formed (or have been actively
540 eroded) very recently.

541 The spatial distribution of the depressions and their fully enclosed shapes suggests that
542 they are not simply scour structures, but have a genetic relationship with debris avalanche

543 deposition. The depressions east of Deposit 1 lie in a region where failure of the pre-
544 existing seafloor sediment occurred during the Deposit 2 landslide [Watt *et al.*, 2012b;
545 Crutchley *et al.*, 2013]. The structures may be collapse pits in younger sediment produced
546 by seafloor subsidence or fluid venting driven by compaction within the underlying
547 landslide deposit.

548

549 **4. Implications for landslide processes**

550 ***4.1. The source and composition of Deposit 1***

551 The rocks exposed in Deposit 1 include near-vent and subaerial lithologies, consistent
552 with English's Crater being the major source of material in the deposit. This correlation
553 places an age of 11.5–14 ka on the formation of English's Crater, which is significantly
554 older than the 6 ka minimum age provided by dates of infilling deposits [Smith *et al.*,
555 2007; Boudon *et al.*, 2007].

556

557 ***4.1.1. Subaerial source region***

558 English's Crater and the Tar River Valley display two volcanic facies [Harford *et al.*,
559 2002]: near-vertical walls of massive lava crop out to the west (Chances Peak; age
560 unknown) and south (Galways Mountain, 112 ka; Perches Dome, 24 ka); and radiating
561 fans of crudely bedded lava breccias (rock fall and block-and-ash flow deposits) crop out
562 at the northern and lower margin of English's Crater and along the Tar River Valley.
563 Block-and-ash flow deposits on the east coast, south of Spanish Point, have radiocarbon
564 ages of 19.7 and 24.0 ka [Roobol and Smith, 1998] and can be traced towards English's
565 Crater. They may be associated with Perches Dome, given their similar age. Similar lava

566 breccias between Chances Peak and Galways Mountain, as well as deposits dated at 16–
567 19 ka on the west side of the island, in Fort Ghaut, suggest elevated levels of extrusive
568 volcanism on Montserrat between 16 and 24 ka. However, the remains of Perches dome
569 are the only exposed Soufrière Hills lavas from this time period. It is possible that a much
570 more extensive lava-dome complex of this age formed the source of the Deposit 1
571 landslide, also removing sections of massive lava from older domes to form the near-
572 vertical cliffs currently exposed around English’s Crater. A relatively deep-seated
573 collapse, centred on the vent region, is supported by the high proportion of
574 hydrothermally altered material in Deposit 1. At least three extensive fumarole and hot
575 spring systems existed inside English’s Crater prior to 1995 (Lang’s, Cow Hill New and
576 Tar River), providing evidence of intense hydrothermal activity in this area [*Roobol and*
577 *Smith, 1998*].

578

579 *4.1.2. Incorporation of submarine material*

580 A single observation of a clast (Figure 2f) with contrasting surfaces of fresh andesite and
581 weathered, tube-worm encrusted andesite, provides the only direct evidence for the
582 incorporation of submarine material within Deposit 1. This conflicts with morphological
583 observations: the maximum plausible subaerial failure volume of $\sim 1 \text{ km}^3$, based on
584 combining the Tar River Valley and English’s Crater depressions, with pre-failure
585 elevations of $>1100 \text{ m}$, is too small to account for the volume of Deposit 1 (1.7 km^3). The
586 chute cut into Montserrat’s eastern flank also suggests that submerged material formed
587 part of the landslide. Such material would likely comprise carbonate and reworked,

588 polymict volcanic clasts. The absence of these lithologies suggests that the surface
589 exposures of Deposit 1 may not be representative of the deposit as a whole.

590 The correlation of Deposit 1 with the large-volume 12–14 ka turbidite east of Montserrat
591 [Trofimovs *et al.*, 2013] (see Section 2.3.1) also implies a submarine component to the
592 event. The turbidite comprises approximately equal proportions of volcanoclastic and
593 bioclastic grains, in contrast to the entirely volcanic lithologies exposed in Deposit 1. If
594 the two events are related, then the bioclastic component of the turbidite must derive
595 from seafloor material disaggregated during landslide emplacement. The shelf chute
596 aligned with Deposit 1 provides supporting evidence of such a process. Given the
597 absence of submarine lithologies within surface exposures of the Deposit 1 hummocks,
598 the submarine component of the landslide may be concentrated disproportionately within
599 the unexposed matrix facies between the debris-avalanche deposit hummocks.

600

601 ***4.2. Emplacement mechanisms and comparison with subaerial debris avalanche*** 602 ***deposits***

603 *4.2.1. Deposit morphologies*

604 Deposit 1 is morphologically and texturally similar to many subaerial debris-avalanche
605 deposits. The rounded hummocks of the deposit, comprising heterogeneous mixtures of
606 deformed and frequently altered monomict domains, are typical of many subaerial
607 examples [e.g. Glicken, 1996; Shea *et al.*, 2008; Clavero *et al.*, 2002]. The fan-shaped
608 morphology of Deposit 1 is comparable to freely spreading deposits such as those at
609 Galunggung and Mombacho volcanoes [Siebert, 1984; Shea *et al.*, 2008], and indicative
610 of granular avalanche emplacement processes [cf. Paguican *et al.*, 2014]. Landslide

611 mobility indices [cf. *Griswold and Iverson, 2008; Iverson et al., 2015*] for Deposit 1 are
612 also within the range of typical values for subaerial volcanic debris avalanches ($L/H = 7$
613 and $A/V^{2/3} = 36$, based on parameters in *Lebas et al. [2011]*) [*Legros, 2002; Griswold and*
614 *Iverson, 2008*].

615 In contrast to Deposit 1, Deposit 2 forms a continuous elongate deposit, and its mobility
616 is at the high end of the range defined by subaerial volcanic debris avalanches ($L/H = 16$
617 and $A/V^{2/3} = 47$, based on parameters in *Watt et al. [2012b]*), which partly reflects the
618 incorporation and secondary failure of large volumes of seafloor-sediment within the
619 deposit [cf. *Watt et al., 2012a,b*]. Deposit 2 has a central thickness of over 100 m, and a
620 surface marked by isolated blocks set within the more continuous landslide mass (as
621 indicated by seismic reflection profiles [*Crutchley et al., 2013*]). Although this mass may
622 be disaggregated and mixed, the blocks are competent, intact fragments of the initial
623 volcanic failure region. They are hundreds of meters across, and have sub-vertical sides
624 that reach over 100 m in height. Observations of the Wembley block and a large block to
625 the south show that they comprise bedded sequences of volcanoclastic breccia, suggestive
626 of marginal and probably near-surface portions of a subaerial lava-dome complex. The
627 blocks result in a prominent morphological front within the thick, central part of Deposit
628 2 [*Watt et al., 2012b*]; the well-exposed southern blocks are closely aligned with the
629 southern lateral margin of the deposit, and the Wembley block lies near the northern
630 margin (Figure 1). The deposit morphology is similar to the Icod debris avalanche
631 deposit, north of Tenerife [*Masson et al., 2002*], which has several kilometer-scale blocks
632 at its lateral margins. *Masson et al. [2002]* conclude that the Icod deposit shape and block
633 distribution is characteristic of coarse-grained debris flow processes [cf. *Major and*

634 *Iverson, 1999*], and suggest that this behavior reflects the high proportion of pyroclastic
635 material in the landslide. Our observations do not show evidence that the Deposit 2
636 failure mass was significantly different to that of Deposit 1, or was rich in friable
637 pyroclastic material, but there is good evidence of extensive seafloor-sediment failure
638 concomitant with the volcanic landslide [cf. *Watt et al., 2012a,b*]. This potentially
639 produced a mixed landslide, with high proportions of fine-grained, clay-rich material.

640

641 *4.2.2. Large-block transport*

642 Hummocks in subaerial debris avalanche deposits are frequently cored by large,
643 deformed blocks of the failure mass [*Crandell et al., 1984; Glicken, 1991; Paguican et*
644 *al., 2014*]. Partial disaggregation, extensional faulting and shearing of these blocks
645 produces the broadly rounded hummock form. The large blocks of Deposit 2 differ from
646 these hummocks in that they have undergone no deformation beyond the initial
647 fragmentation that produced them. The vertical sides, and angular, upright form of the
648 Deposit 2 blocks, as well as their relatively long transport distance, also contrasts with
649 Toreva blocks, which occur in proximal regions of some debris-avalanche deposits and
650 are often rotated, with a morphology that reflects the extensional failure planes of the
651 fragmenting mass [*Siebe et al., 1992; Wadge et al., 1996; Paguican et al., 2014*].

652 The bedded breccias that characterize the Deposit 2 blocks might be expected to
653 disaggregate relatively readily in a debris avalanche. Their preservation as intact
654 fragments of the failure mass may therefore be evidence of an emplacement mechanism
655 that limited block interaction and basal deformation (at least for the small number of
656 outsized blocks near the deposit margins), and may also reflect damping of block

657 collision in the aqueous environment [cf. *de Blasio*, 2013]. Volcaniclastic breccias, as
658 massive and bedded units, also characterize the megablocks in landslide deposits north of
659 Oahu, Hawaii [*Yokose*, 2002], although the failure and transport mechanism is not
660 necessarily similar to that of Deposit 2. Seismic reflection profiles show that the Deposit
661 2 blocks are rooted within a continuous landslide deposit (Figure 12), suggesting that
662 block emplacement is not explained by low-friction transport of individual fragments on
663 a lubricated basal surface of wet sediment (i.e. as characterizes isolated outrunner blocks
664 in some submarine rock avalanches [*de Blasio et al.*, 2006; *de Blasio*, 2013]). Rather, the
665 blocks appear to have been passively rafted within the main landslide mass, without any
666 clear evidence for rotation around a horizontal axis, and pushed towards the margins
667 during continued landslide movement [cf. *Major and Iverson*, 1999]. The lack of
668 subaerial volcanic-debris-avalanche analogues for outsized intact blocks such as those in
669 Deposit 2 may indicate that the development of debris-avalanche masses with sufficient
670 proportions of fine-grained, water-saturated sediment to maintain elevated pore-fluid
671 pressures may be more easily acquired in a submarine environment, via mixing and
672 entrainment of marine sediment.

673

674 **5. Summary and conclusions**

675 This study presents results of the first detailed ROV investigations of multiple submerged
676 landslide deposits around an island-arc volcano. Coupled with other methods of
677 investigation, such as coring, bathymetric mapping and geophysical data, the direct
678 observations offered by ROVs significantly strengthen the interpretation of the sources of

679 material and the processes operating during the emplacement of large landslides around
680 volcanic islands.

681 Our observations indicate that Deposit 1 (1.7 km^3) is similar to many subaerial volcanic
682 debris-avalanche deposits, and is dominated by hydrothermally altered material likely to
683 have originated from a collapse of the near-vent region of the Soufrière Hills volcano.
684 This is surprising, given the large proportion of bioclastic material in a turbidite that
685 correlates stratigraphically with Deposit 1, and a submerged eroded chute associated with
686 the event. However, we infer that the bioclastic component within the turbidite is
687 predominantly derived from pre-existing seafloor sediment disrupted by the emplacement
688 of Deposit 1 and eroded by associated turbidity currents. Our observations suggest that
689 Deposit 1 occurred at 11.5–14 ka through the collapse of altered lava domes erupted at
690 16–24 ka, the relics of which form Perches Dome.

691 A much larger (10 km^3) landslide occurred at ~ 130 ka, forming Deposit 2. Although this
692 deposit was mostly inaccessible to ROV observation, we were able to study a large block
693 of volcanoclastic breccias that represents a single intact fragment of the subaerial volcano.
694 Its petrology is consistent with pre-130 ka Montserrat lavas. The lower part of the block
695 exposes breccia set within a hemipelagic mud matrix, which was most likely acquired
696 through vigorous erosion of pre-existing seafloor sediment during block transport. The
697 intact, oversized blocks within Deposit 2 were rafted within a relatively mobile debris
698 avalanche mass, and are best exposed near the margins of this elongate deposit.

699 Two landslide deposits to the south of Montserrat have very different source lithologies.
700 Deposit 3 is morphologically similar to Deposit 1, but comprises fresher, denser lavas.
701 We infer that it results from a shallower seated collapse, rather than a landslide that cut

702 deeply into a hydrothermally altered edifice. This is consistent with the absence of a
703 prominent source scar for the deposit. Deposit 5 is dominated by blocks of reef rock, and
704 demonstrates that large landslides on the flanks of volcanic islands may occur without
705 involvement of the active volcanic edifice, but can arise from instabilities on the
706 carbonate-dominated shelves that may form around these islands.

707

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721

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937

938 **Figure Captions**

939 **Figure 1**

940 Topographic and bathymetric map of Montserrat, showing the offshore debris-avalanche
941 deposits 1, 2, 3 and 5 [Lebas *et al.*, 2011; Watt *et al.*, 2012b]. Deposits 1, 3 and 5 are well
942 exposed near the seafloor, while Deposit 2 is partially buried but evident from the
943 bathymetric expression of individual large blocks. Dive sites discussed in the text are
944 marked: Isis dives, from cruise JC83, are prefixed I; Hercules dives, from cruise NA037,
945 are prefixed H. Selected vibracore locations, collected on cruise JCR123 [Trofimovs *et*
946 *al.*, 2008, 2010], are also marked. Points prefixed NA037 show the location of samples
947 discussed in the text, and numbered points refer to images in subsequent figures. Isopachs
948 for the 12–14 ka turbidite are taken from Trofimovs *et al.* [2010].

949

950 **Figure 2**

951 ROV images from hummocks within Deposit 1. **(a)** Map of image locations (see Figure
952 1) in this and subsequent figures. Core locations have the prefix JR123, while NA037
953 marks sample locations referred to in the text. **(b)** A dense, shattered lava block in contact
954 with yellow, hydrothermally-altered material and fresh lava breccia along convolute
955 margins. **(c)** Dense lava breccias in contact with hydrothermally-altered red and yellow
956 deformed domains. **(d)** Sheared and stretched deformation within hydrothermally altered
957 domains. **(e)** Vein-like hydrothermal alteration cutting across clast-supported dense lava
958 breccias. **(f)** Lava block with clear division between fresh and colonized surfaces,
959 potentially indicating a submarine origin for some material mobilized in the Deposit 1
960 landslide. **(b, c, d and f)** are from dive H1308 and **e** from I219.)

961

962 **Figure 3**

963 Images from a hummock at the northern edge of Deposit 1 (Figure 2; dive I219). **(a)**
964 Patchy erosion of the hemipelagic cover over Deposit 1, providing a window into the
965 hummock surface and exposures through the overlying sediment. **(b)** Top surface of
966 hummock, showing typical exposure of hydrothermally altered volcanic rock. **(c)**
967 Exposure through the sedimentary drape over Deposit 1, showing a basal layer of dense,
968 gray, angular clasts overlain by a bedded sequence of hemipelagic sediment and mixed
969 volcanoclastic gravel. **(e)** Coarse, hemipelagite-matrix-supported to clast-supported lithic

970 breccias beds overlying the hummock surface. The basal, monomict bed of angular gray
971 lava clasts contrasts with the overlying polymict beds.

972

973 **Figure 4**

974 Images of pale fine-grained sediment, interpreted as hemipelagic mud, within exposures
975 of the Wembley block (locations in Figure 2a). **(a)** Hemipelagic mud-rich interval of the
976 block surface, near the base of the SE side of the block (position on Figure 5), with lava
977 clasts defining crude stratification. **(b)** Mixed volcanoclastic breccia in the upper half of
978 the Wembley block, comprising dense grey, red and black lava clasts (position on Figure
979 5). Beneath a covering of recent grey volcanoclastic sand, pale mud (center) occurs in a
980 small isolated patch, encasing volcanic clasts. **(c)** Typical appearance of pale mud, with a
981 pitted and sculpted surface, in places preserving stretched or sheared fabrics, suggestive
982 of a cohesive, clay rich hemipelagite. **(d)** Crudely bedded polymict, matrix-supported
983 breccia of volcanic clasts embedded in a white to pale hemipelagite mud matrix (outlined
984 in yellow), unconformably overlying a monomict clast-supported lava breccia at a high
985 angle, at the base of the SW side of the block. The right hand panel shows schematic
986 interpretations of the contrasting hemipelagite-rich breccia at different exposures around
987 the Wembley block. In images **a** and **b**, the hemipelagic mud appears to form a matrix to
988 the primary lithology of the block (although how and when this is acquired is open to
989 interpretation – see text), but in image **d** it forms a post-emplacement talus derived from
990 reworked material.

991

992 **Figure 5**

993 A visual log, reconstructed from ROV imagery, of a transect up the exposed surface on
994 the SE side of the Wembley block (map in Figure 2). The surficial exposure may not be
995 representative of internal stratigraphy of the block. A white cohesive material encases
996 volcanic clasts across much of the lower half of the Wembley block, and is interpreted as
997 hemipelagic mud. This material is rare in the upper part of the block. The uppermost part
998 of the block exposes interbedded grey volcanoclastic sands and pale hemipelagic mud,
999 very similar in appearance to material sampled in the JR123 vibrocores from the
1000 surrounding seafloor [Trofimovs *et al.*, 2008, 2010]. Pie charts indicate the relative

1001 proportions of exposed surface area accounted for by different components. Modal and
1002 maximum lithic clast diameters, in centimeters, are given in italics and bold, respectively
1003 (in several cases two modes are apparent).

1004

1005 **Figure 6**

1006 Images of the Wembley block lithologies (locations in Figures 2 and 5). **(a)** Monomict
1007 red and gray lava breccias and massive fresh angular lava blocks. **(b)** Massive single lava
1008 block within side of Wembley block. **(c)** Massive matrix-supported breccia of volcanic
1009 clasts within a white to pale hemipelagite mud matrix. Dark coloration may be due to Fe-
1010 Mn encrustation (arrows). In some cases (lower arrow) the color contrast suggests
1011 variable encrustation in a single clast. **(d)** Succession of two monomict lava breccias
1012 (black overlain by red) in the upper part of the Wembley block, interpreted as block-and-
1013 ash flow deposits. **(e)** Andesite boulder with jig-saw fit fracture implying impact with
1014 nearby blocks during emplacement of the Wembley block. **(f)** Hemipelagic mud bed
1015 exposed at the top of the Wembley block, overlying a recessive bed of volcanoclastic sand
1016 (Figure 8). **(a-c)** are from dive H1308, and **d-f** from I217)

1017

1018 **Figure 7**

1019 Images of carbonate samples (NA037-002 (975 m) and -005 (942 m); Figure 1) from the
1020 Wembley block. **(a)** NA-37-005: a pelagic limestone comprising planktonic foraminifera
1021 (including *Globorotalia*), planktonic gastropods (heteropods and pteropods). Rounded
1022 patches of micrite with few bioclasts are likely burrow fills. Shallow-water benthic
1023 foraminifera are rare. **(b)** Hand specimen of NA037-002, a shallow-water limestone with
1024 rhodoliths of coralline algae and a variety of bioclasts. The same rock is shown in **(c)**,
1025 where rhodoliths have been extensively bored by a clionid sponge, and shallow-water
1026 bioclasts including foraminifera and bivalves are present, along with abundant peloids of
1027 probable microbial origin, and in **(d)**, where large benthic foraminifera (*Amphistegina*,
1028 upper right), coralline algae (center), bivalve fragments (original aragonite replaced by
1029 calcite spar, lower middle), and partially dissolved peneroplid foraminifera (lower left) all
1030 occur within a matrix of micrite and calcite spar. Some volcanic crystals and rock
1031 fragments are also present.

1032

1033 **Figure 8**

1034 Stratigraphic logs and new radiocarbon dates (Table 1) of core JR123-21 taken from the
1035 top of the Wembley block (Figure 2) compared with ROV imagery from the eastern side
1036 of the Wembley block (Figure 5), an exposure through seafloor sediment to the east
1037 (Figure 11c), and stratigraphic logs of cores described in Trofimovs et al. [2013]. Site
1038 locations shown in Figure 2.

1039

1040 **Figure 9**

1041 Images of block exposures in Deposits 3 and 5 (Figure 1; dives H1309 and H1310). **(a)**
1042 Radially-fractured dense lava block with dark Fe-Mn encrustation. This is the dominant
1043 lithology exposed at the surface of Deposit 3. **(b)** Polymict breccias of altered sub-
1044 angular and scoriaceous volcanic clasts, forming a possible surficial deposit overlying
1045 Deposit 3. **(c)** Carbonate cemented conglomerate of rounded lava cobbles (beach type
1046 rock) in Deposit 5. **(d)** Karstic weathering in a reef block in Deposit 5. Field of view ~3
1047 m. **(e)** An overhead view of a weathered carbonate reef block in Deposit 5. **(f)** Slab-like
1048 carbonate blocks within Deposit 5. Similar lithologies were observed on the SW flank of
1049 Montserrat, encrusting the submerged flank of the island.

1050

1051 **Figure 10**

1052 Images of carbonate samples (NA037-025 (806 m) and -026 (823 m); Figure 1) from
1053 Deposit 5. **(a)** NA037-026, a limestone comprising coated rounded and sub-angular
1054 volcanic clasts in a carbonate matrix. In **(b)**, the coating is shown to comprise a mixture
1055 of calcareous algae and other biota, whereas the surrounding matrix contains fine-grained
1056 volcanic material, micrite and calcite spar. A similar matrix and algal-coated grain is
1057 shown in **(c)**, as well as shallow water fossils (e.g. benthic foraminifera *Amphistegina*,
1058 top left). The coating in sample -026 is shown in more detail in **(d)**, where an algae
1059 nodule is encrusted by foraminifera, serpulid and microbial filaments. Sponge spicules
1060 occur in the surrounding matrix. **(e)** NA037-025, a well-sorted bioclastic grainstone,
1061 comprising bioclasts and minor volcanic grains cemented by an isopachous fibrous

1062 calcite fringe (f). Bioclasts include peneroplid foraminifera, coralline algae and bivalve
 1063 fragments.

1064

1065 **Figure 11**

1066 ROV images of circular erosional or collapse structures forming within young seafloor
 1067 sediment (locations in Figures 1 and 2). (a) Shallow dish-like pockmarks in Deposit 5,
 1068 cutting a scarp in hemipelagic sediment all around the margin. (b) Overhead view of a
 1069 relatively deep (~5m) pockmark in Deposit 5. A sharp, circular wall marks positive relief
 1070 beyond the margin of the structure, with a streaking, radiating pattern on the seafloor
 1071 outside the structure. The wall cuts steeply through seafloor strata of interbedded
 1072 hemipelagite and volcanoclastic sand. (c) Pockmark wall beyond the margin of Deposit 1,
 1073 east of the Wembley block, and overlying part of Deposit 2 (Figure 1). This seafloor
 1074 stratigraphy is exposed, showing four distinct hemipelagite layers, present throughout the
 1075 region in cores collected in JR123 [Trofimovs *et al.*, 2008, 2010] (Figure 8). Exposure of
 1076 these young depositional layers suggests recent erosion.

1077

1078 **Figure 12**

1079 A schematic cross-section through the landslide deposits east of Soufrière Hills,
 1080 Montserrat, summarising the main observations made for Deposits 1 and 2 in this study.
 1081 The vertical section and scale are based on seismic profiles through the deposits [cf.
 1082 *Crutchley et al.*, 2013; *Karstens et al.*, 2013].

1083

1084 **Table 1**

1085 Radiocarbon ages of monospecific planktonic foraminifera (*Globigerinoides ruber*)
 1086 picked from hemipelagic mud in core samples constraining the ages of Deposits 1 and 2.

Sample name	Pub. code	Depth below core top (cm)	Conventional age (yr BP) (1 σ error)	Calibrated age range ^c (cal yr BP)	$\delta^{13}\text{C}_{\text{VPDB}}\%$ ± 0.1
JC83-VC1-10 ^a	52752	10–11	1340 (37)	964–781	1.2
JC83-VC1-31 ^a	52753	31–32	3857 (37)	3930–3689	1.6
JC83-VC1-44 ^a	52754	44–45	5615 (37)	6135–5906	1.3
JR123-21-C10	402765	10–11	1870 (30)	1510–1331	0.5
JR123-21-C25	402766	25–26	4760 (30)	5188–4870	1.4
JR123-21-C39	402767	39–40	7450 (30)	7978–7833	1.2
JR123-21-C62	393246	62–65	38940 (400)	43139–42035	0.8
JR123-21-C84	402768	84–85	>43500	NA	0.4
JR123-21-B10	402769	99.5–100.5	>43500	NA	0.1
JR123-21-B22	393247	112–113	30280 (150)	34266–33692	0.6
JR123-21-B71	402770	160.5–161.5	39150 (410)	43311–42141	0.4

JR123-21-B76	402771	166.5	38390 (380)	42763–41710	0.5
JR123-21-B83	393248	173–174	39180 (320)	43191–42263	0.3
JR123-54 ^{a,b}	12994	235	6802 (35)	7406–7294	0.9
JR123-54 ^{a,b}	12995	242	6330 (35)	6895–6685	0.9
JR123-54 ^{a,b}	23055	273	8794 (37)	9525–9395	1.0
JR123-54 ^b	333973	280	8700 (40)	9465–9269	0.1
JR123-54 ^b	333974	284	8600 (40)	9391–9121	1.7
JR123-54 ^b	333975	294.5	9350 (40)	10272–10109	4.4
JR123-54 ^b	333976	303	10830 (50)	12534–12085	1.1

^a Analysed at the NERC Radiocarbon Facility in East Kilbride, UK, following the procedure described in *Trofimovs et al.* [2013]. Publication codes are SUERC- followed by the listed number; All other samples analysed at Beta Analytic Inc. Laboratories, Miami.

^b Ages previously published in *Trofimovs et al.* [2013].

^c Calibrated using OxCal4.2 [*Bronk Ramsey*, 2009] and the Marine13 calibration curve [*Reimer et al.*, 2013]. Calibrated ranges reported at the 95.4% confidence interval. BP refers to years before 1950 A.D.

1087

1088 **Supporting Information**

1089 **Table S1**

1090 Descriptions and locality information for samples discussed in the text.

1091 **Videos S1 to S10**

1092 Selected video clips from ROV dives shown in Figure 1.























