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Environmental crises at the Permian–Triassic mass extinction

Jacopo Dal Corso^{1*}, Haijun Song^{1*}, Sara Callegaro², Daoliang Chu¹, Yadong Sun³, Jason Hilton⁴, Stephen E. Grasby⁵, Michael M. Joachimski⁵, Paul B. Wignall^{6*}

¹State Key Laboratory of Biogeology and Environmental Geology, School of Earth Science, China University of Geosciences, Wuhan, China.

²Centre for Earth Evolution and Dynamics, University of Oslo, Oslo, Norway.

³GeoZentrum Nordbayern, Friedrich-Alexander Universität Erlangen-Nürnberg (FAU), Erlangen, Germany.

⁴School of Geography, Earth and Environmental Sciences, University of Birmingham, Birmingham, UK.

⁵Geological Survey of Canada, Natural Resources Canada, Calgary, Alberta, Canada.

⁶School of Earth and Environment, University of Leeds, Leeds, UK.

*Correspondence should be addressed to J. Dal Corso (j.dalcorso@cug.edu.cn), Haijun Song (haijunsong@cug.edu.cn), and P.B. Wignall (P.B.Wignall@leeds.ac.uk).

Key points

- The Permian–Triassic mass extinction (252 Ma) resulted in a substantial reduction of global biodiversity, with the extinction of 81–94% of marine species and 70% of terrestrial vertebrate families.
- Sedimentary, palaeontological and geochemical records during the mass extinction indicate that a cascade of environmental changes caused the extinction.
- The environmental changes can be linked (and attributed to) the effects of volcanic emissions (for example CO₂, SO₂, and metals) during the eruption of the Siberian Traps large igneous province.
- Inferred volcanically driven environmental perturbations include global warming, oceanic anoxia, oceanic acidification, and (potentially) ozone reduction, acid rain, and metal poisoning.
- The crisis on land likely started ~60–370 kyrs before that in the ocean, indicating different response times of terrestrial and marine ecosystems to Siberian eruptions.

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- The causes of marine extinctions are inferred from geochemical and sedimentary evidence, but the reasons for the earlier terrestrial ecological crises remain poorly understood, but likely include rapid atmospheric change.

34 **Abstract**

35 The link between the Permian–Triassic mass extinction (PTME; 252 Ma) and the emplacement of the
36 Siberian Traps Large Igneous Province (STLIP) was first proposed over 30 years ago. However, the
37 complex cascade of volcanic-driven environmental and biological events that led to the largest known
38 extinction in life’s history is still difficult to reconstruct. In this Review, we critically evaluate the
39 geologic evidence and discuss the current hypotheses surrounding PTME kill mechanisms. Data
40 indicate that the initial STLIP extrusive/pyroclastic volcanism was coeval with widespread crisis of
41 terrestrial biota and marine animal species stress at high northern latitudes. The following onset of
42 extensive magmatic intrusions is linked with the rapid (~60 kyr) extinction of 81–94% of marine
43 species. The terrestrial to deep water extinctions are thought to have been caused by a combination of
44 global environmental perturbations driven by the emissions from STLIP. Nevertheless, it remains
45 difficult to understand the ultimate reason for the exceptional severity of the PTME. Future research
46 needs improved geochronology of STLIP and sedimentary sequences (especially terrestrial) to better
47 resolve the timing of volcanic phases and extinctions. Further ecological and physiological studies are
48 needed to understand temporal and spatial extinction patterns. Improved modelling is necessary to
49 reconstruct the causal relations between volcanism, environmental perturbations and ecosystem
50 collapse.

51

52 **Introduction**

53 Many mass extinction [G] events punctuated the history of life and changed evolutionary trajectories
54 ¹. Most of past biological crises are coeval with the emplacement of Large Igneous Provinces (LIPs)
55 [G], which drove widespread environmental perturbations. LIP emissions of CO₂ and other gasses are
56 comparable to current anthropogenic emissions, and future climate projections predict a scenario
57 similar to the major Phanerozoic extinctions. Hence, understanding past events will help define the
58 tipping points that lead to a major biological crisis ².

59 The Permian–Triassic mass extinction (PTME; 252 Ma) was the most severe biological crisis of the
60 Phanerozoic (Fig. 1). It almost completely eliminated Palaeozoic fauna and flora, setting the stage for
61 the evolution of modern ecosystems. Across the Permian–Triassic boundary (PTB), 81–94% of
62 marine species went extinct ^{3–5} (Fig. 2 and 3). On land, 89% of tetrapod genera and 49% of families
63 disappeared ⁶ (Fig. 4). Recovery began in the Early Triassic ^{7–9}, but became significant only in the
64 Middle Triassic, five million years later ^{10–12}.

65 Data from the fossil, sedimentary, and geochemical record of the PTME suggest there were major
66 environmental changes in marine and terrestrial settings^{13–15} (Fig. 2 and 3). The global crisis is
67 coeval with the emplacement of the Siberian Traps Large Igneous Province (STLIP)^{16–18} (Fig. 5),
68 that saw a relatively rapid (<1 Ma) eruption of 2–7 million km³^{19–22} of basalt, together with volcanic
69 emissions of CO₂, SO₂, halogens and metals that were capable of causing global climate and
70 environmental catastrophe.

71 Detailed timing of events has improved remarkably in recent years thanks to advances in high-
72 precision radioisotope dating [G], and high-resolution biostratigraphy [G] and chemostratigraphy [G]
73 studies (especially C-isotope and Hg stratigraphy; BOX 1 and 2). Analysis of events from 252 million
74 years ago at a high temporal resolution allowed identification of distinct phases of STLIP eruptions
75^{18,23} and separate pulses of extinction among marine animals^{4,24,25}. Particularly interesting
76 developments include the increasing evidence that the terrestrial crisis was very likely underway
77 several tens to hundreds of thousands of years before the marine extinction^{26–28}, clearly indicating
78 that the PTME was not a single, instantaneous catastrophic event. Whilst these findings are expanding
79 knowledge of STLIP volcanism, environmental changes, and extinction patterns, linking them
80 remains difficult. The geological record tells a complex and partly obscure story of multiple, co-
81 occurring phenomena, all playing a role in perturbing the ecosystems, and all probably interlinked in a
82 cascade of environmental disasters.

83 In this Review we discuss the PTME pattern in the ocean and on land, the age and volcanic style of
84 the STLIP, the evidence of a link between STLIP phases and the PTME, and the environmental crises
85 triggered by the volcanic emissions and their role in the extinction and observed selectivity. We
86 discuss the apparent diachrony between some recorded environmental changes and extinctions. We
87 then construct a likely chronology of the events based on the current evidence, propose a working
88 hypothesis for future research, and highlight the open problems.

89

90 **Pattern of the PTME**

91 The exact temporal relationship between the marine and terrestrial extinctions is still debated.
92 However, there is increasing evidence for an earlier onset of the terrestrial crisis and marine stress at
93 high northern latitudes. The age and pattern of the marine PTME at low latitudes are very well
94 constrained, and these provide a stratigraphic framework that allows the level of terrestrial crisis to be
95 pinpointed. Here we examine the marine crisis first.

96 **Marine extinction.** Across the PTB, the Palaeozoic evolutionary fauna [G] was devastated at all
97 ecological levels, resulting in the largest marine extinction of the entire Phanerozoic (Fig. 1a and
98 Supplementary Information). It has been estimated that 81–94% of marine species went extinct^{3–5}.

99 The PTME appears to have been selective (Fig. 3). Some groups completely disappeared, such as
100 trilobites, rugose and tabulate corals, fusulinid foraminifers, and blastoid echinoderms^{4,29–31}; others,
101 such as rhynchonelliforms (articulate brachiopods), crinoids, stenolaemate bryozoans, calcisponges,
102 radiolarians, ammonoids, and ostracods, came close to annihilation with only a handful of surviving
103 species^{4,32}; whilst a few groups, including bivalves, gastropods, conodonts, and fishes, experienced
104 “only” severe to moderate extinction rates^{4,33–35} (Fig. 2). Extinction selectivity is not only evident in
105 the taxonomic composition of the marine fauna, but also in ecological and physiological traits. Body-
106 size selectivity is seen in foraminifers, conodonts brachiopods, and bivalves^{36–39}, with larger bodied
107 organisms showing higher extinction rates, but this factor is less obvious in other groups^{34,36,40} (Fig.
108 3).

109 It has been shown that physiologically buffered taxa that can regulate intracellular chemistry and
110 counterbalance environmental chemical changes, like molluscs, ostracods, arthropods, and fish,
111 experienced lower extinction rates than unbuffered taxa such as brachiopods and echinoderms^{41–44}
112 (Fig. 3). Moreover, non-motile taxa could in general be affected more by changing environmental
113 conditions than motile animals, especially swimming animals¹²: Fish were relatively little affected
114 compared to other groups⁴⁴. However, statistical analysis shows that selectivity between these two
115 groups was not significant (Fig. 3).

116 Taxa with limited geographic distribution are generally more prone to extinction than widely
117 distributed groups because they are more dependent on local environmental conditions. However, this
118 phenomenon is not so strong during the PTME^{43,45,46} (Fig. 3), suggesting harsh marine environments
119 were global in extent; there was no escape even for cosmopolitan species. Indeed, weak geographic
120 range selectivity appears to be a general pattern with the major mass extinctions, being observed also
121 at the end-Triassic and the end-Cretaceous⁴⁵ (Fig. 1a).

122 The pace of the PTME extinction pattern is long debated³², with contrasting hypotheses of gradual vs
123 abrupt extinction, and single vs discrete extinction pulses. The gradual disappearance of marine
124 species observed in several PTB successions below the main extinction horizon^{3,47} could be
125 attributed, for most groups, to the Signor-Lipps effect [G] in the fossil record whereby rarer species
126 are last recorded some time before their final demise^{3,48,49}. The one exception is the ammonoid

127 extinction pattern observed in the succession of Iran where, considering the age confidence interval of
128 each species, ammonoid diversity indeed shows a gradual decline before the PTME⁴⁷.

129 The marine extinction was a geologically brief event in the latest Permian–earliest Triassic⁵⁰. Well-
130 studied sections from South China, such as Meishan GSSP [G], show enormous losses at the base of
131 the latest Permian *Clarkina meishanensis* zone^{3,51}. Abrupt extinction in the latest Permian has also
132 been documented in Italian foraminifera and Iranian brachiopods, foraminifera, and algae^{48,52}. Thus,
133 the crisis is often referred to as the end-Permian mass extinction, implying a single, geologically
134 instantaneous (~30 kyr) event at the end of the Permian, just before the stratigraphic PTB^{3,51} (Fig. 2).
135 However, a significant diversity of Permian taxa, including brachiopods, foraminifers and ostracods
136 survived beyond this level to become extinct either in the final part of the Permian or in the earliest
137 Triassic^{53–55}. Combining data from sections encompassing a spectrum of water depths shows that
138 there was an especially intense final extinction phase in South China at the base of the *Isarcicella*
139 *isarcica* Zone⁴. It is noteworthy that the second extinction pulse was proportionally intense but not as
140 large in magnitude in terms of species loss.

141 Multi-phase extinction pattern has also been reported in the Dolomites, Italy^{24,25}. Here, the main/first
142 phase (which correlates to the first extinction pulse in South China), witnessed the loss of calcareous
143 algae, foraminifera, and large-sized brachiopods and molluscs, with a genus extinction rate of 64%²⁵
144 (Fig. 2). 68% of survivors and newcomers subsequently went extinct in the interval ranging from the
145 upper *H. changxingensis* to lower *H. parvus* zones²⁵, with a second extinction pulse that is somewhat
146 slightly earlier than the second pulse in South China⁴ (Fig. 2).

147 Adequate sampling obviously play an important role in the assessment of the extinction pattern. When
148 fossil occurrences are rare, statistical confidence in the precise timing and number of pulses of
149 extinction declines. Moreover, the importance of examining extinction patterns in different
150 environments is highlighted by the record of foraminifer which shows a single extinction pulse in
151 shallow platform facies but two episodes of extinction in deep slope facies⁵⁶. In contrast, brachiopods
152 suffered two episodic extinctions in shallow platform⁵⁷ and deep slope³ and basin facies⁵⁸.

153 The latest U-Pb zircon ages from Meishan show that the two pulses of extinction happened at 251.941
154 and 251.880 Ma, respectively, separated by an interval of ~60 kyrs⁵⁹ (Fig. 2). It is unclear if the
155 intervening interval should be considered a time of high stress or if the two pulses of extinction
156 represent discrete events. The interlude was certainly an intriguing time, it saw the proliferation of
157 microbialites and oolitic strata in low latitudes^{60,61}, whilst origination rates [G] increased⁴, pointing

158 to a temporary improvement in environmental conditions either within the extinction interval, or
159 between the two main extinction pulses (Fig. 2).

160 **Terrestrial extinction.** The PTME was the Phanerozoic's largest and most severe extinction of
161 terrestrial plants and animals, at all latitudes and trophic levels (Fig. 4). Terrestrial floras suffered a
162 worldwide catastrophic die-off of many plant groups in different geographical and climate zones
163 during the PTME, which reset plant evolutionary history and was followed by an Early–Middle
164 Triassic “coal gap”: an interval in which peat-forming communities disappeared^{10,26,62–67}. Some have
165 argued that plant losses were much more moderate compared to animals^{68,69}, but the unprecedented
166 abrupt shutdown of peat formation is a clear signal for major loss of terrestrial biomass across the
167 PTB.

168 Plant fossil records from South China show that diversity and abundance of the tropical rainforest-
169 type *Gigantopteris* flora experienced a sharp decrease with a loss of 95% of species and 50% of
170 genera^{70,71}. The Permian gymnosperm-dominated floras of North China and Russia experienced
171 similar catastrophic losses at the same time⁷². In the southern hemisphere (Gondwana), the coal-
172 forming *Glossopteris* flora went extinct (and coals disappeared) to be replaced by seed-fern shrubs
173 (*Dicroidium*)^{10,67,73}, a stratigraphically long-ranging genus that survived the PTME and migrated
174 polewards from low-latitudes⁷⁴. The subsequent earliest Triassic floras throughout Eurasia and the
175 southern continents were dominated by lycopods, especially *Pleuromeia*.

176 During the crisis, palynological data show widespread spore abundance spikes, accompanied by high
177 abundances of spore tetrads and teratological [G] pollen, evidencing stressed conditions^{67,75–83} (Fig.
178 4). Peak abundances of *Reduviasporonites* have been noted as marking fungal or algal bloom events.
179 The former attribution would indicate proliferation of fungal saprotrophs during terrestrial ecosystem
180 collapse^{67,84–86}, but the affinity of *Reduviasporonites* and its ecological significance remain
181 controversial^{87,88}.

182 Major changes at the base of the terrestrial food-web, for example in the structure of floral
183 communities from luxurious forests to less productive lycophyte-dominated floras, triggered a
184 cascade of extinction in terrestrial ecosystems at all higher trophic levels^{81,89}. It is therefore
185 noteworthy, but perhaps not surprising, that the mass extinction was the only one to severely affect
186 insects, with losses of 30% of orders and 50% of families^{90–92}.

187 Tetrapods were also severely impacted by the PTME with numerous families lost. Complex latest
188 Permian ecosystems, dominated by herbivorous pareiasaurs, dicynodonts and carnivorous
189 gorgonopsians, were replaced by ones with archosaurs and synsids^{93,94}. Global tetrapod generic

190 data suggest there was an 89% generic loss of tetrapods near the PTB ⁶. Such losses within tetrapods
191 could have happened during a “sustained extinction interval” of up to ~1 Myr, as seen in the fossil
192 record from the Karoo Basin (South Africa) ⁹⁵ (Fig. 4), although it is difficult to determine rates in the
193 low quality tetrapod record. Many niches disappeared with studies showing the loss of all small fish-
194 and insect-eaters, medium and large herbivores and large carnivores in Russia ^{94,96,97}. Coupled to the
195 terrestrial extinction and vegetation loss, fluvial style changed across the PTME from meandering to
196 braided rivers and aeolian systems, as observed in the successions of the Karoo Basin, Russia,
197 Australia and North China (for example ref. ⁹⁸).

198 The timing of terrestrial ecosystem crisis relative to extinctions in the oceans is debated. Some studies
199 argued that the terrestrial extinction was coeval with that in the oceans based on radioisotopic dating,
200 chemostratigraphy and terrestrial information recorded in marine deposits ^{71,99,100}. However, recent
201 high-resolution studies show that the terrestrial ecosystems were already stressed before the marine
202 PTME (Fig. 4). Evidence of an earlier terrestrial crisis is based on improved stratigraphic frameworks
203 including C-isotope stratigraphy (BOX 1), Hg (and Ni) spikes (BOX 2), magnetostratigraphy, and
204 high-precision dating methods ^{26–28,95,101–103} that supersede previous lower resolution studies. This
205 earlier crisis is seen in palaeofloras from the high-latitude Sydney Basin ²⁶, in the tropical peatland
206 ecosystems in equatorial South China ²⁷, in the flora and fauna of North China ¹⁰², and in the tetrapod
207 losses in the high-latitude Karoo Basin ^{28,95} (Fig. 4). Recurrent wildfire and abnormal pollen in the
208 latest Permian also indicate that terrestrial ecosystems were under great stress before their collapse
209 and prior to the subsequent marine extinction ^{26,27,78,101,102,104}.

210 It is important to note that the terrestrial fossil record primarily derives from lowland settings,
211 especially lacustrine and riparian environments where much sediment accumulates. This is especially
212 the case for the plant record which is dominated by plants from wetlands, due to their good
213 preservational conditions in such settings ¹⁰⁵. Much less is known about plants from drier and upland
214 habitats which rarely fossilize, although evidence of upland vegetation loss during the PTME has been
215 inferred from changing sedimentary facies in the Karoo Basin and Russia ¹⁰⁶. Wetland extinctions,
216 that eliminated the *Glossopteris* and *Gigantopteris* mire communities, were not mirrored by equally
217 severe losses in drier habitats dominated by conifers and pteridosperms ^{107,108}. This pattern may partly
218 explain why the palynological record, which includes data of widely dispersed pollen and spores from
219 drier and upland habitats mixed with those of lowland floras, often shows a much more muted
220 extinction intensity during the PTME ^{107,108}.

222 **The trigger**

223 The two main agents suggested to be responsible for the environmental changes that led to the
224 PTME—which will be discussed in the next section—are extraterrestrial impact or large-scale
225 volcanism.

226 Extraterrestrial impact as the trigger of the PTME was proposed on the basis of geochemical (He^3 in
227 fullerenes) and sedimentological (for example, the finding of chondritic meteorite fragments in the
228 sedimentary record) data, and of the interpretation of a large structure found in the Indian ocean, the
229 Bedout High, as a purported impact crater of supposed PTME age^{109,110}. However, the
230 extraterrestrial impact hypothesis has been largely rejected because data have been proven to be
231 difficult to reproduce, and the age and interpretation of the impact structure and geochemical changes
232 have been questioned^{111,112}. Most scientists now agree that there is negligible evidence to support
233 extraterrestrial impact as the cause of the PTME.

234 In contrast, overwhelming data support that the PTME was triggered by the eruption of the STLIP.
235 The STLIP was emplaced during the Permian–Triassic transition in the continental Tunguska Basin
236¹⁸, the adjacent West Siberian Basin¹¹³ and Taimyr Peninsula^{21,114}, with a poorly-constrained,
237 original volume between 2 and 7 million km^3 ^{19,21,22} (Fig. 5). Changes in geochemistry and mode of
238 emplacement of the STLIP magmas during its history, as seen in the most accessible lava pile sections
239 (Norilsk, Putorana and Maymecha-Kotui) and in outcrops and boreholes from the Tunguska, Taimyr
240 and West Siberian basins^{18,115–119}, reveal three phases of magmatic activity²³.

241 The oldest lava flows and pyroclastic [G] deposits were alkaline [G], mafic to ultramafic, and high in
242 TiO_2 (>2 wt%)¹¹⁷. These magmas have a deep, pyroxenitic mantle source [G], and were rich in
243 magmatic Cl^{119–122}. The gases released by this high-volume, initial phase of magmatism were
244 probably dominantly mantle-derived, products of recycled oceanic crust entrained by the mantle
245 plume^{119,122}, although additional crustal sources are also likely^{120,121,123}. This first phase of
246 magmatism started just before 252.27 ± 0.1 Ma, with extensive effusive activity taking place over the
247 ~300 kyr preceding the marine PTME^{18,23}. Intriguingly these early eruptions did not produce global
248 changes in the $\delta^{13}\text{C}$ record nor in Hg concentrations (Fig. 5). However, northern latitude marine
249 records downwind of the eruption site show pre-extinction shifts in the $\delta^{13}\text{C}$ and Hg records^{124–126},
250 suggesting limited atmospheric mixing of volatiles released in this early eruption phase. This first
251 STLIP phase appears to have been coeval with widespread terrestrial crisis^{26,27} (Fig. 5).

252 The second, tholeiitic [G] phase consisting of sills [G] and intrusions in the Tunguska Basin and
253 Taimyr Peninsula began at 251.907 ± 0.067 Ma, and is coeval with the onset of the marine PTME (Fig.

254 5). The apparent absence of effusive and/or explosive activity during this phase is debated and may
255 reflect a lack of sampling in the region ^{114,127–131}. The tholeiitic phase magmas were derived from a
256 shallow mantle pyroxenitic-peridotitic source ^{119,132}, that underwent widespread interaction with the
257 crust ^{132–136}, and is poor in juvenile volatiles [G] ¹³⁷. During this subvolcanic phase the STLIP
258 intruded a succession of coal, shales, sandstones, evaporites and carbonates in the Tunguska Basin
259 ^{125,129,130} and their baking may have liberated a large amount of both greenhouse gases and
260 halocarbons ^{123,129,130,132,138,139}. Explosive basalt pipes and breccia diatremes are widespread in the
261 Tunguska Basin and are interpreted to have been the result of this gas generation and violent escape to
262 the atmosphere ^{128,129,140,141} (although at least some may have erupted later in the Triassic ¹³¹).
263 Contact metamorphism of organic-carbon rich sediments around large-scaled sill intrusions in Taimyr
264 and the Tunguska Basin (Siberia) likely produced large quantities of isotopically light CO₂ and CH₄
265 capable of changing the C-isotope signature of the atmosphere and ocean ^{21,125,129,138}. Indeed, the
266 onset of intrusive magmatism, given analytical uncertainty, coincides with the start of the negative
267 shift in δ¹³C in the *C. yini* Zone (251.999 ± 0.039 Ma; Fig. 2 and 5) ⁵⁹, providing indirect support for
268 this notion.

269 The third and last STLIP phase started at 251.483 ± 0.088 Ma with renewed lava extrusion (alkaline),
270 and ongoing intrusive activity (alkaline and tholeiitic), in the Maymecha-Kotui 118 and Taimyr
271 regions ²¹. These magmas are interpreted as extremely deep and hot products of a volatile-rich source
272 ¹⁴², and were likely enriched in mantle-derived CO₂ ¹²¹. A maximum age for the end of this phase,
273 and STLIP activity overall, is placed at 250.2 ± 0.3 Ma ¹⁸. A gradual recovery of the δ¹³C curve (BOX
274 1) towards pre-extinction levels is observed in the sedimentary record during the third phase (Fig. 5).

275

276 **Environmental crises**

277 The consequences of the volcanic emissions from the STLIP are considered in this section, including
278 the emissions produced by contact metamorphism caused by magmatic intrusions in the host rocks,
279 mainly CO₂ and CH₄, SO₂, halogens (for example, Cl, F, Br and halocarbons) and metals (for
280 example, Hg, Cu). The multiple effects of these emissions are considered separately and are likely to
281 have operated at different stages in the history of the PTME crisis (Fig. 6). Inferred environmental
282 crises include global warming, oceanic anoxia, oceanic acidification, and (potentially) ozone
283 reduction, acid rain, and metal poisoning.

284 **Global warming.** Among the gases released by volcanism, SO₂ has the potential to trigger short-term
285 cooling episodes over a duration only a little longer than the eruptive interval ^{143,144}. This is too short

286 an interval to be detectable, given temporal resolution possible in deep time. Furthermore, it is a moot
287 point whether such brief cooling intervals are capable of causing appreciable environmental stress,
288 although episodic cooling events set in a context of longer-term warming could have damaged the
289 ecosystems¹⁴⁴.

290 Instead, global warming, due to CO₂ and CH₄ emissions is the clearest signal to emerge from the
291 eruption of LIPs. Temperatures across the PTB have been reconstructed using oxygen isotopes ($\delta^{18}\text{O}$)
292 in conodont apatite and brachiopod calcite. Conodont $\delta^{18}\text{O}$ data from low latitude sections from Iran
293^{145,146}, Armenia¹⁴⁷ and South China^{148–152}, all indicate significant low-latitude warming of 8–10°
294 C from the latest Permian to Early Triassic (Fig. 2 and 4). Despite differences between the analysed
295 localities due to different palaeolatitude or depositional settings, calculated sea surface temperatures
296 (SST) indicate pre-extinction (*C. nodosa*/*C. yini* Zone) SSTs of ~24–30°C that rapidly increased
297 across the PTB and into the earliest Triassic SSTs (*C. isarcica* Zone), ultimately peaking at ~35–39°C
298¹⁴⁷. Conodont $\delta^{18}\text{O}$ records indicate that temperatures increased over an interval of ~39 kyrs¹⁴⁷,
299 although curiously the warming slightly postdates the initial shift in carbonate $\delta^{13}\text{C}$ ^{147,151} (Fig. 2).

300 Warming of 8–10° C likely resulted in a loss of performance of many marine organisms. High
301 ambient temperatures increase metabolic activity and enhance oxygen demand, causing limited
302 functional capacity of oxygen supply culminating in hypoxemia, anaerobic metabolism and loss of
303 protein function¹⁵³. Thermal tolerance of marine organisms is also linked with an organisms' level of
304 metabolic activity, deoxygenation and also oceanic acidification^{154,155}. Thus, warming may have
305 been a major agent of the mass extinction. However, the first phase of the PTME occurred at the onset
306 of warming when conditions may still have been relatively amenable (Fig. 2), and warming may have
307 had a stronger role in the second extinction pulse in the earliest Triassic⁴. Lethally hot temperature
308 may have induced selective extinction of marine animals and poleward migration^{148,156,157}.

309 However, two opposite patterns of selective extinction across latitudes had been reported^{156,157} one
310 showing the highest extinction rate in the high latitudes¹⁵⁷, the other documenting the highest rate in
311 the tropics¹⁵⁶. This discrepancy is likely due to the different statistical schemes used. Whilst higher
312 polar extinction has been inferred, the study only considered the end-Permian extinction pulse¹⁵⁷
313 whilst higher tropical extinction is calculated considering two pulses (end-Permian and earliest
314 Triassic)¹⁵⁶. Likewise, in the marine fossil record of South China⁴, the calculated extinction rate is
315 57% if only the first pulse is taken into account, and 93% including both pulses.

316 Poleward migration of about 10–15° is also observed in tetrapods^{148,158}. In plants, elevated
317 temperatures and droughts can inhibit photosynthesis, increase photooxidative stress due to higher

318 irradiance, burn leaves, and limit plants' growth and yield, and ultimately cause their death ¹⁵⁹.
319 Warming could have also increased the prevalence of wildfire by increasing seasonality and drought
320 (Fig. 3), for example as proposed for the records of South China where high charcoal abundance is
321 found in strata recording the ~60 kyrs initial decline of $\delta^{13}\text{C}$, up to the onset of the marine crisis
322 ^{27,104,160} (Fig. 2). Elevated fire activity would have aided post-fire run-off and erosion ¹⁰⁴.

323 However, the terrestrial extinction appears to have started before the warming trend inferred from the
324 $\delta^{18}\text{O}$ of conodont apatite (Fig. 2 and 4). In South China, declining $\delta^{13}\text{C}$ values coincide with high
325 charcoal abundance ²⁷ (Fig. 3) suggesting atmospheric $p\text{CO}_2$ was increasing during the interval of
326 higher wildfire activity. Also, along the northwestern margin of Pangea marine environmental stress
327 began prior to the main extinction event, suggesting that higher latitude oceans were deteriorating as
328 the terrestrial extinction initiated ^{125,161,162}. Curiously, these changes occurred prior to the warming
329 trend recorded by conodont $\delta^{18}\text{O}$ data.

330 **Oceanic anoxia.** The PTB coincides with a eustatic sea-level rise and the development of an oceanic
331 anoxic event (OAE) [G] that has been directly implicated as a cause of the crisis ¹⁶³. However, marine
332 anoxia during transgression is often encountered in the geological record, raising the question of why
333 these conditions caused such a severe extinction crisis? There are likely to have been three reasons:
334 the anoxia extended in some regions into extremely shallow waters ¹⁶⁴, although oxic refugia
335 remained ¹⁶⁵; the Panthalassa superocean also become anoxic throughout much of the water column
336 ^{111,166}; the OAE persisted, with varying intensities, for several million years into the Middle Triassic,
337 prolonging the stressful conditions for marine life ¹⁶⁷⁻¹⁶⁹. Thus, both the extent and duration of anoxia
338 were exceptional by Phanerozoic standards.

339 Evidence for anoxia is diverse and found in a broad range of environments. Organic-rich, pyritic,
340 black shales, the typical manifestation of anoxic deposition, are best developed in the deep ocean
341 sections now found in the accreted terranes in Japan and New Zealand ^{111,166,170}. Black shales are less
342 common in shelf and epicontinental seaways, especially in tropical settings, perhaps due to high
343 organic matter remineralization rates in hot sea water. In the low-latitude carbonate setting of Tethys,
344 anoxic facies are usually developed as laminated, pyritic micrites such as in northern Italy ⁵⁰. In
345 northern Boreal shelf seas, anoxic facies include finely-laminated, argillaceous strata and pyritic
346 sandstones with abundant framboidal pyrite [G] ^{161,164}.

347 Intensity of marine anoxia and its extent are inferred from geochemical data. The uranium isotope
348 ratio of $^{238}\text{U}/^{235}\text{U}$ recorded in limestones shows a shift to lower values immediately prior to the first
349 phase of mass extinction (Fig. 2): a change attributed to the accelerated removal of ^{238}U in anoxic

350 bottom waters¹⁷¹. The degree of anoxia driven metal drawdown was such that the oceans become
351 depleted in trace metals¹⁷². The scale of anoxia also affected the ocean's sulphate budget.
352 Increasingly heavy sulphate-sulphur isotope values in the Early Triassic, relates to removal of
353 isotopically light pyrite sulphur, suggesting reduced seawater sulphate concentrations¹⁷³. Biomarkers
354 also provide evidence for oxygen-poor conditions including the presence of isorenieratane, an
355 indicator that anoxic conditions extended into the photic zone¹⁷⁴.

356 The development of intensive anoxia profoundly altered the oceans' nutrient structure. Phosphorus
357 recycling enhances under anoxic conditions¹⁷⁵ and, when combined with higher continental runoff,
358 this leads to high phosphorus availability in the water column. However, nitrogen rather than
359 phosphorus was more likely the limiting nutrient in the anoxic oceans of the time^{176,177}. Thus,
360 nitrogen isotope ratios show a significant decrease, from values up to ~10‰ to ~0‰, in most of the
361 basins across the PTB^{177,178}. This suggests strong denitrification accompanied the onset of global
362 anoxia, likely due to a fundamental shift from a nitrate-dominated to an ammonium-dominated
363 nutrient supply which would normally favour nitrogen-fixing diazotrophs. However, diazotrophs
364 require molybdenum and iron for nitrogen fixation and yet these are efficiently removed from anoxic
365 waters, thereby causing a decrease in the ocean's total fixed-nitrogen and low levels of productivity
366¹⁷⁶. Some alternative scenarios favour productivity increase during the extinction interval, driven by
367 enhanced nutrient run-off¹⁷⁹, but these fail to account for the micronutrient limitations of diazotrophs
368 in euxinic waters, as well as the absence of organic-rich shales in the Early Triassic¹⁸⁰.

369 The ultimate cause of the Permian–Triassic OAE has long been attributed to the effects of STLIP with
370 warming and more sluggish ocean circulation usually invoked^{111,163}. The Community Earth System
371 Model with its embedded biogeochemical cycles, shows that an 11°C sea-surface temperature rise (a
372 realistic value supported by $\delta^{18}\text{O}$ evidence¹⁴⁸; Fig. 2 and 5), combined with increased freshwater
373 runoff into high latitude seas, greatly increases ocean stratification and decreases meridional overturn
374 circulation¹⁵⁷. The result is a dramatic decrease in seafloor oxygenation. The model also successfully
375 replicates regional variations with the best ventilated area shown to be the Perigondwanan margin of
376 southern Tethys¹⁵⁷, a finding that closely matches field evidence from this region⁵⁴.

377 **Oceanic acidification.** Another potentially harmful effect of massive CO₂ injection into the
378 atmosphere–ocean system is oceanic acidification (Fig. 6). Huge amounts of CO₂ entering the oceans
379 acidifies water and decreases carbonate saturation. Evidence for oceanic acidification at the PTB
380 comes from boron isotope ($\delta^{11}\text{B}$) and calcium isotope ($\delta^{44/40}\text{Ca}$) records^{43,181,182} (Fig. 2), and the
381 sediment record^{43,162,183}. However, data from $\delta^{11}\text{B}$ of bulk carbonates, used to signify acidification
382 during the second phase of the PTME during the *I. isarcica* Zone¹⁸⁴, are now generally considered

383 not to actually reflect marine pH¹⁸⁵. Instead, a composite $\delta^{11}\text{B}$ record from pristine brachiopod shells
384 from the Southern Alps (Italy) and South China shows a decline in $\delta^{11}\text{B}$ values, which suggests
385 lowering of seawater pH, between the onset of the negative C-isotope excursion and the base of the
386 *parvus* Zone, just above the PTB¹⁸² (Fig. 2). This composite $\delta^{11}\text{B}$ record needs, however, further
387 validation in other sections and improvement of temporal resolution. Ooidal limestones are
388 widespread during the inferred lower pH interval²⁵, indicating saturated conditions, and the analysed
389 brachiopods come from interbedded levels of microbialites¹⁸². The prevailing carbonates suggest that
390 under saturation was not achieved. Acidification could have happened in very brief pulses, which are
391 not recorded by low-resolution datasets, rather as a relatively longer event between the onset of the
392 marine extinction and the earliest Triassic¹⁸².

393 A negative $\delta^{44/40}\text{Ca}$ shift during the PTME interval has been linked to the injection of CO_2 from the
394 STLIP activity on the basis of its stratigraphic correlation with the negative $\delta^{13}\text{C}$ excursion¹⁸⁶.
395 Instead of solely indicating oceanic acidification, Ca-isotope data modelling suggests that a complex
396 scenario controlled seawater $\delta^{44/40}\text{Ca}$ changes, involving CO_2 release, acidification, reduced skeletal
397 carbonate sink, enhanced weathering of shelf carbonates, changes in carbonate mineralogy and
398 changes in seawater saturation state^{187,188}. In detail though, the negative $\delta^{13}\text{C}$ excursion (in bed 24 at
399 Meishan) predates the negative $\delta^{44/40}\text{Ca}$ shift (which occurs above bed 25¹⁸⁶; Fig. 2), complicating the
400 interpretation of the relationships between the Ca- and C-isotope records. Similar negative $\delta^{44/40}\text{Ca}$
401 excursions, recorded by both conodont apatite and bulk carbonate, are seen at the same stratigraphic
402 interval in other localities^{188–191}.

403 More indirect evidence for oceanic acidification comes from the fossil record which shows that the
404 crisis saw the preferential extinction of physiologically unbuffered taxa, with low metabolisms and
405 high energy demand for the production of calcium carbonate skeletons (for example corals,
406 brachiopods, calcareous sponges, and foraminifera), whilst well-buffered taxa (for example bivalves,
407 gastropods, ammonoids and conodonts) could have survived the crisis relatively better^{4,43,192}.

408 Analysis of the microstructure of brachiopod shells provides evidence to suggest a role for
409 acidification in brachiopod extinction losses. All brachiopod groups suffered severe losses with the
410 diverse Strophomenata going extinct. The Rhynchonellata fared somewhat better and it has been
411 suggested that their higher shell organic content enabled them to better survive acidified conditions
412¹⁸³. However, at lower taxonomic order the Rhynchonellata suffered severe losses and their story
413 during the PTME could also be described as a successful re-radiation of the survivors in the earliest
414 Triassic that saw some genera become widespread¹⁹³. In addition, the preferential extinction of

415 coarsely ornamented ammonoids supports the pressure of oceanic acidification on shell-building costs
416 for shelled animals ¹⁹⁴.

417 Along the north western margin of Pangea there is also a gradual loss of carbonate producers through
418 the late Permian creating an empty ecologic niche that was filled by siliceous sponges expanding from
419 deep environments to become the dominant organism in late Permian shallow shelves ^{161,162},
420 suggesting decreasing pH prior to the extinction.

421 **Ozone disruption.** High abundance of teratological sporomorphs during the PTME (Fig. 4) has been
422 attributed to increased UV-B radiation due to disruption of the ozone layer ^{76,78,81,195}. Experiments on
423 living *Pinus mugo* showed increasing exposure of plants to UV-B radiation induced malformation on
424 pollen grains similar to those observed at the PTME and, although all trees survived, their fertility
425 markedly decreased ⁸¹. Therefore, higher UV-B radiation and lower plant fertility may have triggered
426 a collapse of the whole terrestrial ecosystem by shutting down most primary productivity.

427 Ozone depletion could have been driven by the release of halogens and halocarbon compounds from
428 volcanic activity and the combustion of coals and evaporites intruded by STLIP ^{196–198} (Fig. 5).
429 However, the first explosive phase of STLIP activity appears to be coincident with the early terrestrial
430 decline of plants and the first occurrences of teratological sporomorphs (Fig. 4), whilst the release of
431 halocarbons (for example CH₃Cl) from contact metamorphism (intrusive phase) is thought to have the
432 strongest impact on the ozone layer ^{197,198}. Teratological sporomorphs are found throughout the
433 PTME (Fig. 4), but ozone is quickly (~10 yrs) restored in the atmosphere, hence making a long-term
434 disruption of the global ozone unlikely ¹⁹⁸.

435 **Acid rain.** Teratological sporomorphs (Fig. 4) alone are not a direct evidence of UV-B radiation, as
436 they could be the result of other stresses such as acid rain ^{81,198} and metal poisoning ⁸². Acid
437 deposition can potentially kill plants, phytoplankton, vertebrates and invertebrates in terrestrial
438 aquatic ecosystems, and acidification of non-calcareous soil results in leaching of important nutrients
439 (Ca, Mg and K), with the effect of weakening plants and increasing their mortality rate ¹⁹⁹.

440 Magmatic degassing of SO₂ and halogens from STLIP could have driven acid rain ¹⁹⁸. Earth system
441 modelling shows that, alongside the previously discussed ozone damage, S injected into the
442 stratosphere during STLIP pyroclastic activity (Fig. 5) could have triggered extensive acid rains at the
443 PTB, although these were only severe (pH = 2) in the Northern Hemisphere ¹⁹⁸.

444 Possible direct evidence of acid rain comes from one section in northern Italy, where the abundance of
445 vanillin—a product of pH-dependent enzymatic decomposition of organic matter in soil—could

446 suggest pulses of soil acidification²⁰⁰. Vanillin peaks occur before the marine extinction interval
447 (latest Permian)²⁵. Hence, acid rains may have affected terrestrial ecosystems already before the
448 onset of the marine extinction (Fig. 5). Significantly, PTB palaeosols in Antarctica show high
449 chemical weathering but no indication of acid conditions; there was no leaching of Ca and Mg²⁰¹.

450 Other geochemical evidence for acid rain comes from sulfur isotope and concentration records in the
451 Karoo Basin (South Africa), where higher accumulation of sulfide was interpreted as the effect of
452 high sulfate supply to the freshwater environment from acid rain²⁰². However, the terrestrial
453 extinction in the Karoo Basin began before the S geochemical changes, making their significance
454 moot. Currently, except for these local datasets, there is no conclusive evidence that widespread acid
455 rain triggered the terrestrial collapse in the latest Permian, especially not in the southern hemisphere.

456 ***Metal poisoning.*** Potentially, metal poisoning may have occurred in marine environments, where an
457 increase of concentration of toxic metals (Hg, Cr, As, and Co) is observed^{125,126,161}. High
458 concentrations of Hg, the most toxic metal, may have been reached after the marine extinction, when
459 the reduction of bioproductivity could have led to a decrease of Hg drawdown by organic matter and
460 its potential build-up in marine environments to toxic levels, before it was removed by sulphide
461 deposition¹²⁶.

462 A coincidence between a peak of teratological lycophte spore tetrads [G] and high Hg and Cu
463 concentrations has been found a short distance above the terrestrial extinction level in South China,
464 indicating that the survivor plants might have experienced stress caused by higher metal
465 concentrations in the environment⁸² (Fig. 4). Reduced plant transpiration, changes to the hydrological
466 cycle and climatic drying following terrestrial vegetation loss may have resulted in reduced water
467 availability in freshwater ecosystems leading to such metal concentrations increase. Hg is generally
468 thought to derive from volcanic activity²⁰³, but Hg isotopes and modelling of Hg cycling indicate
469 that, superimposed on a general increase of volcanic Hg deposition across the PTME, further Hg
470 could have been released into the environment due to massive oxidation of terrestrial organic matter
471 and soil following the collapse of land ecosystems²⁰⁴ (BOX 2). Similar behaviour could have sourced
472 Cu⁸². Hence, the increase of metal loading in South China during the PTME might actually be the
473 consequence of the demise of the *Gigantopteris* rainforests and wetland species²⁰⁴.

474

475 **Linking kill mechanisms and extinction patterns**

476 The latest high-resolution chronology of the PTME (Fig. 6) suggests that the terrestrial ecological
477 disturbance could have started 60–370 kyr before the marine extinction^{26–28,95,101–103}. This was
478 coincidental with the initial, mostly explosive phase of STLIP. However, the temporal resolution of
479 the terrestrial extinctions remains more poorly known than that of the marine extinctions and may
480 have been spread over ~1 Myr⁹⁵.

481 The terrestrial extinction mechanism is not clear, and mainly inferred by indirect, often local, and
482 mainly palaeontological proxies (Fig. 6). Increased seasonality during the initial stage of the negative
483 $\delta^{13}\text{C}$ (Fig. 6) could have led to increase of wildfires²⁷. Declining $\delta^{13}\text{C}$ values coeval with higher
484 charcoal abundance suggest addition of isotopically light CO_2 to the ocean–atmosphere system and
485 that warming may have played a role. However, the available temperature proxy from marine settings
486 (conodont $\delta^{18}\text{O}$) suggests temperatures did not begin to increase until after the terrestrial crisis had
487 begun (Fig. 2 and 5).

488 Temporal decoupling of terrestrial extinctions predating marine extinctions is intriguing and suggests
489 spatial heterogeneity in the extinction patterns and potentially mechanisms. Delayed onset of marine
490 extinctions may be partially related to thermal inertia of the oceans and their higher thermal capacity
491 compared to land that heats and cools quicker²⁰⁵, but ocean turnover times occur in the order of 1000
492 years²⁰⁶ so are unlikely to have operated at a 60–370 kyrs time scale.

493 Terrestrial stress may have come from emissions of SO_2 and halogens and their consequent acid rains
494²⁰², and disrupted ozone shield^{76,81}. Increasing UV-B radiation on Earth's surface and acid
495 depositions could have had lethal effects on terrestrial ecosystems, causing stress to the vegetation,
496 lowering plants' fertility and eventually leading to their death, with repercussions at all higher trophic
497 levels. However, long-term disruption of the global ozone during the PTME is unlikely¹⁹⁸.

498 It is not clear what was the effect on marine ecosystems of the first phase of the STLIP activity (Fig.
499 6). Beds of coal ash and associated Hg spikes are observed in northwest Pangea prior to the main
500 negative $\delta^{13}\text{C}$ excursion as well as Ni isotope anomalies that may record this initial phase of eruptions
501 impacting the terrestrial environment¹²⁵. This region also shows early marine stress^{161,162}, while
502 more equatorial records show no marine impacts.

503 The marine extinction interval has a clear, temporal link with the second mostly intrusive phase of the
504 STLIP and gas emissions, and persisted for <100 kyrs straddling the PTB. There were two pulses of
505 extinction intensity at the beginning and end of this interval although significant losses were also
506 occurring in the interlude interval too.

507 Taxonomic, morphologic, and ecologic selectivity (Fig. 3) and the magnitude of marine extinction
508 suggest that a combination of global warming, anoxia, and oceanic acidification best explains the
509 marine PTME (Fig. 6). Groups intolerant to hypoxia and high temperature were preferentially
510 eliminated during the PTME, suggesting that these stressors played an important role in the extinction
511 of marine animals^{157,207} (Fig. 3). Physiologically buffered taxa experienced lower extinction rates
512 than unbuffered taxa^{42–44} (Fig. 3). Oceanic acidification could have been an important stressor for
513 shelled animals⁴³, as also supported by the preferential extinction of coarsely ornamented ammonoids
514¹⁹⁴ (Fig. 3). Survival animals migrated to higher latitudes or deep seawaters, possibly to escape the
515 hot temperature in equatorial regions or surface seawaters^{56,148,156} (Fig. 6).

516

517 **Summary and future directions**

518 The link between the PTME and the eruption of the STLIP has been well established since the late
519 1990s¹⁷. Dramatically improved absolute dating has strengthened the link to the point where
520 scenarios involving distinct stages of the emplacement history can be linked with consequent
521 environmental changes (Fig. 6). The effects of the eruptions were likely experienced first in terrestrial
522 settings, where plant productivity crashed and coal ceased to form, and in high-latitudes marine
523 settings in the northern hemisphere. The initial explosive phase of the STLIP emplacement may have
524 driven this crisis, including increased seasonality, ozone depletion, with higher UV-B radiation, and
525 acid rain.

526 The marine mass extinction is coeval with the mainly intrusive phase of the STLIP. Increasing fossil
527 and geochemical data resolution indicates that the marine mass extinction could have happened either
528 in two distinct pulses or gradually within an interval straddling the PTB. The thermogenic gases
529 produced by the interaction of magma with the intruded sediments introduced into the PTB
530 atmosphere–ocean system triggered a rapid temperature rise, a decline in ocean ventilation, and ocean
531 acidification, which led to the mass extinction. However, despite the large amount of available data
532 and significantly improved geochronology, the reconstruction of the complex co-occurring phenomena
533 interlinked in the fatal cascade that drove the PTME remains difficult.

534 Future research direction should aim at improving the spatial and temporal resolution of datasets from
535 PTME terrestrial records. High-precision U-Pb dating of ash beds and detrital zircons, together with
536 magnetostratigraphy and chemostratigraphy, will increase the chronological constraints of the
537 terrestrial crisis, clarifying the delay between the beginning of the extinction on land and in the ocean.

538 Improved spatial coverage of high-precision stratigraphic syntheses will further evaluate extinction
539 pattern heterogeneity.

540 Detailed evaluation of PTME palynological assemblages will give a more comprehensive picture of
541 through-ranging taxa to understand dynamics and composition of upland “refugial” or survivor floras.
542 The occurrence of teratologies in sporomorphs must be studied in different plants groups, at different
543 latitudes and throughout the PTME, to identify their ultimate cause and understand whether it
544 interested worldwide flora, and at which stages of the event. Further S-isotope and biomarker analysis
545 of PTME terrestrial successions could strengthen the evidence of acid rains during the terrestrial
546 extinction interval.

547 The temporal relationship between warming and extinction, both on land and on the ocean, remains
548 problematic, and further studies, including modelling, should try to understand the apparent lags
549 between the C-isotope, O-isotope, and fossil records. Future high-resolution studies ($\delta^{18}\text{O}$ from
550 conodont apatite or brachiopod shells) will be pivotal in detecting brief temperature changes on the
551 already manifest long-term CO_2 -driven warming trend. However, the current limitation is not the
552 precision of $\delta^{18}\text{O}$ analysis but sample availability. Higher resolution can only be achieved by
553 decreasing the size of conodont samples taken in the field followed by SIMS analyses of individual
554 conodont elements.

555 Further ecological and physiological studies are required to link environmental changes and extinction
556 patterns. Quantitative predictions for extinction selectivity under different changing environmental
557 conditions are needed to distinguish among potential killing stressors. More consistent geochemical
558 ($\delta^{11}\text{B}$) and palaeontological records of ocean saturation are necessary to properly investigate the role
559 of ocean acidification.

560 Furthermore, future endeavours from the geochronology community should be focused on shedding
561 more light on the temporal correlations between the intrusive and effusive realms of the STLIP, which
562 are still weak. Moreover, since most of the STLIP deposits are covered, it is difficult to fully assess
563 the true nature of the eruption history. Drilling programs could significantly expand the knowledge on
564 the history of the STLIP emplacement. The voluminous tephra deposits and the explosive pipes are
565 tangible proof of explosive activity of the STLIP and of gas discharge to the atmosphere. Clarifying
566 the origin and timing of emplacement of these products and structures would contribute greatly to
567 understanding the link between STLIP emplacement stages and global environmental changes.

568 Perhaps the most overriding question from the study of mass extinctions driven by volcanic
569 emissions, of which the PTME is the key example, is what can it tells us about future climate trends.
570 Clearly, extreme global warming can lead to severe consequences for the life but if these effects lie

571 tens of thousands of years in the future, then they are of no geopolitical concern. If changes occur
572 over decades or centuries then their significance increases. Despite the great advances in resolving the
573 details of the PTME, future studies of the crisis should attempt to decipher rates of change on 100–
574 1000 year scale.

575

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577

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1195 AUTHOR CONTRIBUTIONS

1196 JDC coordinated the developing of the article. All authors contributed to the writing of the manuscript
1197 and building of the figures.

1198 COMPETING INTERESTS

1199 The authors declare no competing interests.

1200 **DATA AVAILABILITY STATEMENT**

1201 Data from the Paleobiology Database used for the new calculation of the marine extinction rate are
1202 available in the Supplementary Materials.

1203 **FIGURE CAPTIONS**

1204 **Figure 1. The Permian–Triassic mass extinction and its world.** The PTME, also known as the
1205 “Great Dying”, is the largest extinction of the entire Phanerozoic, with severe losses both in marine
1206 and terrestrial ecosystems. The PTME world consisted in one single continent (Pangea) surrounded by
1207 a vast ocean (Panthalassa), and a giant gulf (Palaeo- and Neo-Thetys). The Siberian Traps Large
1208 Igneous Province erupted 2–7 million km³ of basalt in the northern hemisphere during the PTME. The
1209 biological crisis was the result of the environmental changes triggered by the volcanic emissions from
1210 the Siberian Traps, including the emissions produced by contact metamorphism caused by magmatic
1211 intrusions in the host rocks, such as CO₂, CH₄, SO₂, halogens and metals, into the Permo–Triassic
1212 atmosphere–ocean system. **a)** Newly calculated Gap-filler (GF) extinction rates²⁰⁸ (Supplementary
1213 Note 1) for marine animals show the PTME stands out as the most severe extinction event compared to
1214 other intervals. Along with the PTME, the Ordovician–Silurian, Frasnian–Famennian, end-Triassic,
1215 end-Cretaceous mass extinctions are usually regarded as the largest extinction events of the
1216 Phanerozoic, also known as the “Big 5”. **b)** Palaeogeographic reconstruction of Earth during the
1217 Permo–Triassic transition. Palaeogeography is from ref.²⁰⁹. GSSP = Global Stratotype Section and
1218 Point.

1219 **Figure 2. Marine mass extinction.** Pattern of the extinction in marine settings and major recorded
1220 geochemical changes. High-resolution geochemical data coupled to species richness of different
1221 marine groups. Palaeontological data show two extinction pulses spanning the Permian–Triassic
1222 boundary (PTB). While the first pulse appears to be synchronous in different areas, the second major
1223 pulse of extinction may have been diachronous. Geochemical changes mark the marine extinction
1224 interval (the interval between the two pulses), and testify for major environmental changes coeval to
1225 the biological crisis, as global warming ($\delta^{18}\text{O}$), oceanic anoxia (Uranium-isotope, $\delta^{238}\text{U}$, and sulfur-
1226 isotope of carbonate-associated sulphate, $\delta^{34}\text{S}_{\text{CAS}}$), and ocean acidification (Boron-isotope, $\delta^{11}\text{B}$,
1227 calcium-isotope, $\delta^{44/40}\text{Ca}$). Carbon-isotope ($\delta^{13}\text{C}$) data come from the most updated compilations of
1228 ref.^{210,211}. Oxygen-isotope ($\delta^{18}\text{O}$) data from conodont apatite are from StabisoDB (Stable Isotope
1229 Database for Earth System Research)²¹². $\delta^{18}\text{O}$ data measured with SIMS (Secondary Ion Mass
1230 Spectrometer) have been corrected by a factor of -0.6‰ according to estimates by ref.²¹³ of the offset
1231 between conodont *in-situ* SIMS and bulk IRMS (Isotope Ratio Mass Spectrometer) analyses.
1232 Uranium-isotope ($\delta^{238}\text{U}$) data are from ref.²¹⁰. $\delta^{34}\text{S}_{\text{CAS}}$ data are from ref.^{173,179,214}. $\delta^{11}\text{B}$ data from
1233 brachiopod calcite are from ref.¹⁸². $\delta^{44/40}\text{Ca}$ data are from ref.¹⁸⁶. Species richness from numerous
1234 PTB sections in South China is from ref.⁴. Genera richness from the Dolomites (Southern Alps, Italy)
1235 is from refs.^{24,25}.

1236 **Fig. 3. Extinction selectivity during the Permian–Triassic mass extinction.** The pattern of the
1237 PTME suggests statistically significant extinction selectivity between different ecological groups,
1238 shedding lights on the causes of the marine mass extinction. However, even if selectivity clearly
1239 played a role, high extinction rates are recorded for all marine ecological groups. **a)** Summary of
1240 extinction selectivity trends observed in marine animals: Based on refs. ^{4,37,38,41–44,183,194}. **b)** Extinction
1241 magnitude among different ecological groups in South China ⁴. There are significant differences
1242 (Mann Whitney test, $p < 0.05$) between extinction severity among different ecologic groups, i.e.,
1243 nekton *vs* benthos, buffered *vs* unbuffered, bivalve *vs* brachiopod. Selectivity between motile and non-
1244 motile animals appears to have been less significant ($p = 0.05$). Bars represent 95% confidence
1245 intervals. **c)** Ecological selectivity of global extinctions during the PTME ⁴³. A zero log-odds value
1246 means there is no association between the ecological traits and extinction. The unbuffered and
1247 carbonate-shell genera were preferentially killed ($p < 0.05$). Selectivity between narrow-geographic-
1248 and cosmopolitan-range genera is weaker. Selectivity among genera with calcite shell, infaunal, and
1249 lower abundance of individuals is not significant ($p > 0.05$).

1250 **Figure 4. Terrestrial mass extinction.** Pattern of the extinction in terrestrial settings and major
1251 recorded geochemical changes. Organic C-isotope ($\delta^{13}\text{C}_{\text{TOC}}$), Hg and Hg/TOC, and main biological
1252 events from reference sections of the terrestrial PTME in Northwestern China ^{102,215–217}, South China
1253 ^{27,71,82}, Sydney Basin ^{26,67} and Karoo Basin ^{28,95}, and correlation with the marine $\delta^{13}\text{C}_{\text{carbonate}}$ and
1254 $\delta^{13}\text{C}_{\text{TOC}}$ (Meishan only) records ²¹⁷. The $\delta^{13}\text{C}_{\text{TOC}}$ record from Meishan has been used as a
1255 chemostratigraphic tool to correlate the marine GSSP to the terrestrial sections of Northwestern China
1256 ²¹⁷. Data from the terrestrial PTME records with high-resolution chemostratigraphic data and/or
1257 redioisotopic ages, which allow correlation with the marine PTME, indicate that the terrestrial crisis
1258 started before the marine mass extinction. *Gigantopteris* and *Glossopteris* forests collapsed 60 kyrs
1259 (South China) ^{27,204} to 370 kyrs (Sydney Basin) ^{26,67} before the marine extinction. Wildfire activity
1260 widespreadly increased ^{27,67}. Tetrapods experienced high extinction rates, probably over a relatively
1261 long interval of up to ~ 1 Myr ⁹⁵. Existing data strongly support that the terrestrial crisis started before
1262 the global marine mass extinction. Evidence of also an earlier marine crisis come from high-latitude
1263 northwestern margin of Pangea.

1264 **Figure 5. Link between Siberian Traps, extinction, C-cycle changes and global warming.**
1265 Radiometric ages of the volcanic products (lava, tuff, and sills) of the Siberian Traps Large Igneous
1266 Province (STLIP) and sedimentary Hg geochemistry (BOX 2) indicate this LIP was active during the
1267 PTME, and was linked to injection of isotopically-light carbon into the Permian–Triassic atmosphere–
1268 ocean system, as inferred by the C-isotopes ($\delta^{13}\text{C}$) record (BOX 1), which raised $p\text{CO}_2$ and increased
1269 global temperature, as shown by O-isotopes ($\delta^{18}\text{O}$) of conodont apatite. Different volcanic phases can

1270 be defined: a first mainly pyroclastic phase (lava and tuff), a second mainly intrusive phase (sills), and
1271 a final extrusive phase. **a)** Schematic map of the STLIP (adapted from refs. ^{119,129}) showing the
1272 predominance of lava, pyroclastic and subvolcanic magmatic products over cratonic and non-cratonic
1273 regions of this vast province. M-K stands for Maymecha-Kotuy. **b)** Geochemical data linking the
1274 STLIP to extinction and environmental changes. Dating U/Pb ages of intrusive and extrusive rocks of
1275 the Siberian Traps are from ref. ^{18,21}. Hg and Hg/TOC data are from ref. ^{218,219}. Only Hg/TOC data
1276 with TOC>0.2% have been plotted following the approach of ref. ²⁰³. Source of $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ data as
1277 in Fig. 2. Time span of marine and terrestrial extinction intervals are as defined in Fig. 2. The eruption
1278 of STLIP was very likely the trigger of the Permian–Triassic mass extinction.

1279 **Figure 6. Extinction mechanisms.** Summary of the volcanically-triggered extinction mechanism
1280 inferred from the geochemical, sedimentary and palaeontological record of the PTME and their
1281 recorded effects on biota. The initial mainly extrusive–pyroclastic volcanic phase is coeval to the
1282 initial terrestrial crisis, whilst the onset of intrusive volcanism is coeval with the marine extinction
1283 interval. The different volcanic styles and linked injection of greenhouse gases, halogens and metals
1284 (only the most relevant volcanic/volcanogenic gases are represented in the figure) in the end Permian–
1285 earliest Triassic atmosphere–land–ocean system, triggered a cascade of environmental disturbances,
1286 as summarized in the figure, which firstly affected terrestrial ecosystems and high-latitude marine
1287 environments, and then marine biota. The effects of the environmental changes on the physiology and
1288 ecology of terrestrial and marine biota were multiple, showing that a fatal combination of factors,
1289 sustained for a relatively long interval and each having selective effects on biota, led to the most
1290 severe extinction of the Phanerozoic.

1291

1292 **BOX 1: The Permian–Triassic boundary C-isotope record**

1293 Carbon isotopes ($\delta^{13}\text{C}$) are used as a chemostratigraphic tool to correlate marine and terrestrial
1294 successions around the world. Many high resolution $\delta^{13}\text{C}$ records have been collected across the PTB
1295 and they provide a powerful correlation tool, which helps identifying the PTME interval and link
1296 biological and environmental phenomena recorded in different locations. Major $\delta^{13}\text{C}$ shifts in
1297 carbonate ($\delta^{13}\text{C}_{\text{carbonate}}$)^{220–223}, and marine and terrestrial total organic carbon ($\delta^{13}\text{C}_{\text{TOC}}$)^{221,224,225}, are
1298 documented during the PTME (Figs. 2 and 3). A 3–6‰ negative $\delta^{13}\text{C}_{\text{carbonate}}$ excursion begins
1299 gradually in the lower *C. yini* Zone (*C. bachmanni* Zone) ~60 kyrs below the onset of the marine
1300 crisis^{59,221}, before accelerating to reach a minimum values in the earliest Triassic (*H. parvus* to early *I.*
1301 *isostichia* Zone). Similar shifts are recorded by $\delta^{13}\text{C}$ values from total organic matter, wood and
1302 leaves, allowing correlation of non-marine to marine records (for example refs.^{217,225–228}; Fig. 2 and
1303 3). However, because $\delta^{13}\text{C}_{\text{TOC}}$ is dependent on variable contributions of algal vs. bacterial and marine
1304 vs. terrigenous organic matter, some records display non-parallel trends in $\delta^{13}\text{C}_{\text{carbonate}}$ and $\delta^{13}\text{C}_{\text{TOC}}$, as
1305 documented, for example, at Meishan GSSP section²²⁵. A variety of mechanisms were suggested to
1306 explain the negative $\delta^{13}\text{C}$ shifts by the addition of isotopically light carbon to the exogenic carbon-
1307 cycle reservoirs. Besides soil erosion, reduced primary productivity and destabilization of gas
1308 hydrates, Siberian Traps volcanism and related processes were favoured as the ultimate cause.
1309 Identification of the source of the isotopically light carbon and its $\delta^{13}\text{C}$ signature is critical to estimate
1310 the amount of carbon transferred into the PTB atmosphere–ocean system, and to model atmospheric
1311 $p\text{CO}_2$ increase, temperature rise and seawater pH decline.

1312

1313

1314

1315 **BOX 2: Tracing Siberian Traps activity in the sedimentary record**

1316 Significant increases in mercury (Hg) concentrations above background occur at marine and terrestrial
1317 PTME boundaries globally, and have been attributed to Hg emissions from the Siberian Traps Large
1318 Igneous Province (STLIP)^{27,126,203}. If correct, Hg serves as a ‘fingerprint’ of STLIP in the
1319 sedimentary record, allowing temporal correlation between the eruption and the extinction with
1320 resolution on a millennial time scale²²⁹. As a volatile gas Hg has sufficient atmospheric residence
1321 time for inter-hemispheric mixing, until eventually being transferred through wet or dry deposition to
1322 the marine and terrestrial environment, and after going through various biogeochemical cycling,
1323 eventual geologic sequestration in sediments^{203,229}. In theory then, enhanced Hg emissions related to
1324 the STLIP should be recorded as an Hg spike in sediments²²⁹. This is not definitive though as
1325 concurrent changes in sequestration pathways, such as enhanced bioproductivity and consequent
1326 increased organic matter drawdown, could also create Hg spikes. Careful analyses of Hg data and
1327 sequestration pathways is required before a linkage with STLIP is possible. Stable isotope data (Fig.
1328 3), particularly mass independent fractionation (MIF), support Hg anomalies in offshore marine
1329 deposits being largely derived from a volcanic source²⁰³. However, these same data show nuances in
1330 the Hg cycle. Nearshore deposits have Hg spikes with a MIF signature of terrestrial vegetation²⁰³,
1331 likely related to devastation of forest and swamp ecosystems at that time^{203,204}. Whether Hg
1332 anomalies are directly from volcanos, or indirectly from terrestrial reservoirs released through STLIP
1333 induced global warming, they both serve as a fingerprint (or LIP mark) of STLIP. Resolving the
1334 relative Hg pathways requires further work, along with understanding of how terrestrial and marine
1335 Hg records can be used to resolve the apparent diachronous extinction. Figure is adapted from ref.²⁰³

1336

1337 **GLOSSARY (in alphabetic order)**

1338

1339 ALKALINE

1340 Any rock of a magmatic series presenting a high content of alkali (Na_2O and K_2O) relative to silica
1341 (SiO_2).

1342

1343 BIOSTRATIGRAPHY

1344 Technique to determine the relative age of sedimentary rocks using their fossil content.

1345

1346 CHEMOSTRATIGRAPHY

1347 The study of geochemical variations in sedimentary rocks; Globally-recorded chemostratigraphic
1348 changes are used to correlate sedimentary sequences.

1349

1350 CONODONT

1351 The hard part of an extinct jawless vertebrates, similar to an eel.

1352

1353 EVOLUTIONARY FAUNA

1354 A fauna type that typically shows an increase in biodiversity following a logistic curve, i.e., Cambrian
1355 fauna, Paleozoic fauna, and Modern fauna.

1356

1357 FRAMBOIDAL PYRITE

1358 Aggregates of pyrite (sulfide mineral, FeS_2) with a “raspberry” (“la framboise” in french) aspect. It is
1359 used as a palaeo-redox proxy.

1360

1361 GSSP

1362 Global Stratotype Section and Point. Reference stratigraphic section and level where boundaries
1363 between geological stages, for example between the Permian and the Triassic, are defined.

1364

1365 JUVENILE VOLATILE

1366 A gas species that is dissolved in, or exsolved from, a magma, and is thus newly introduced to the
1367 atmosphere when the magma reaches the Earth's surface.

1368

1369 LARGE IGNEOUS PROVINCE

1370 Rapidly emplaced (<1–5 Myrs) volcanic provinces with areal extents >0.1 million km^2 and volumes
1371 >0.1 million km^3 .

1372

1373 MASS EXTINCTION

1374 Global biological events of greatly elevated extinction rates.

1375

1376 OCEANIC ANOXIC EVENT

1377 Interval of severely reduced dissolved oxygen content in the ocean.

1378

1379 ORIGINATION RATES

1380 The ratio of the number of newly occurring species/genera to the total number over a given geological
1381 period.

1382

1383 PYROCLASTIC

1384 Volcanic rock composed by fragmented pieces of lava. Coarser pyroclastic fragments accumulate in
1385 proximity to the erupting vent, while finer particles can travel hundreds of kilometres.

1386

1387 PYROXENITIC MANTLE SOURCE

1388 A mantle source dominated by the presence of pyroxene and by paucity or lack of olivine. They
1389 represent enriched and very fertile mantle lithologies.

1390

1391 RADIOISOTOPE DATING

1392 Technique to determine the absolute age of rocks using radioactive decay.

1393

1394 SIGNOR-LIPPS EFFECT

1395 A paleontological principle which states that the fossil record of organisms is never complete.
1396
1397 SILL
1398 A tabular subvolcanic magma-body, emplaced roughly concordant or to the general bedding
1399 (stratification or layering) of its host-rocks.
1400
1401 SPORE TETRAD
1402 Four connected immature spore grains in tetrahedral or tetragonal fashion produced by meiotic
1403 microsporogenesis.
1404
1405 TERATOLOGICAL SPOROMORPHS
1406 Pollen and spores that present congenital abnormalities, such as lack of full development and
1407 malformations in their structure.
1408
1409 THOLEIITIC
1410 Sub-alkaline series of magmatic rocks, which undergo iron enrichment during differentiation due to
1411 their poorly oxidised state. Tholeiites are the products of extensive melting of the mantle.
1412
1413
1414
1415
1416

1417